



Spherulite formation in obsidian lavas in the Aeolian Islands, Italy

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ABSTRACT

Spherulites in obsidian lavas of Lipari and Vulcano (Italy) are characterised by spatial, textural and geochemical variations, formed by different processes. Spherulites vary in size from <1 mm to 8 mm, are spherical to elongate in shape, and show variable radial interiors. Spherulites occur individually or in deformation bands, and some are surrounded by clear haloes and brown rims. Spherulites typically contain cristobalite (α , β) and orthoclase, and rhyolitic glass, and grew over an average period of 5 days, with modification at lower temperatures. Heterogeneity relates to formation processes of spherulite 'types' at different stages of cooling and emplacement. Distinct populations concentrate within deformation structures, with variations in shape and internal structure. Crystal Size Distribution (CSD) plots show differing size populations and growth periods. Spherulites which formed at high temperatures show elongation, where deformation triggered further spherulite nucleation and growth. Spherulites formed at mid-glass transition temperatures are spherical, and spherulites are modified at vapour-phase temperatures. Enhanced undercooling, deformation, and modification are therefore pivotal in the development spherulite heterogeneity in obsidian lavas.

Keywords: Spherulites; Obsidian; Aeolian Islands; Glass Transition; Lipari; Vulcano.

INTRODUCTION

Rhyolitic lava flows and domes on the islands of Lipari and Vulcano (Aeolian Islands, Italy; Figure 1) provide key information on the formation and modification of spherulites in obsidian lavas. This is in part due to their widespread occurrence, excellent exposure, and notable textural and geochemical variations. Despite their common occurrence in obsidian lavas worldwide, questions remain regarding the conditions and processes of spherulite formation. For instance, Watkins et al. (2008) highlight the open questions: on what timescale do spherulites form? When do spherulites begin to form? Do spherulites grow below the glass transition temperature bracket? Clay et al. (2013) further stress the need for a detailed evaluation of the link between spherulite formation and deformation

of obsidian lava flows. The preservation of spherulite textures in Aeolian Islands obsidian lavas presents a unique opportunity to classify spherulite heterogeneity, and identify processes responsible for their formation and variation.

Spherulites are radial fibrous crystallites, usually comprised of silica polymorphs and feldspar (Lofgren, 1971a; Swanson, 1977; McArthur et al., 1998; Gardner et al., 2012; Breitzkreuz, 2013). Spherulites are considered to progressively nucleate as a down temperature continuum during syn- and post-emplacement cooling (Clay et al., 2013), originally nucleating in response to an enhanced rate of undercooling brought on by degassing (Kirkpatrick, 1975; Swanson, 1977; Castro et al., 2009; Clay et al., 2013). As well as undercooling, plastic deformation



Figure 1. The Aeolian Islands, South Tyrrhenian Sea (Italy), with the islands of Lipari and Vulcano in the centre of the archipelago, straddling the Tindari-Letojanni fault system (Ventura, 2013).

following extrusion has also been suggested as a trigger for the formation of spherulites (Clay et al., 2013), and is particularly emphasised in organic and inorganic compounds and polymers (Shtukenberg et al., 2012). It is also suggested that spherulites form as a hydration and devitrification texture at lower temperatures (Lofgren 1971b; Swanson et al., 1989). It has been suggested that spherulite growth can occur above, within, and below the glass transition (T_g - 750-600 °C), whereby the material changes from liquid-like to solid via a viscoelastic phase transition (Ryan and Sammis, 1981; Manley, 1992; Davis and McPhie, 1996; Watkins et al., 2008; Gardner et al., 2012; Clay et al., 2013), meaning that emplacement mechanisms of obsidian lavas may play a key role in spherulite formation.

We report here a data set covering twelve obsidian lava flows and domes across Lipari and Vulcano, comprising 5405 individual spherulites, with the focus on the characterisation of their spatial distribution, textural features, and geochemistry. Spherulites have been classified in terms of their physical characteristics, chemical composition, and formation origin. Petrographic observations, mineral and glass geochemistry, laser Raman spectroscopy, X-ray diffraction (XRD) and Crystal Size Distribution (CSD) methods have been used in order to constrain spherulites formation in obsidian lavas in the Aeolian Islands obsidian suite. These observations and variations provide an indication of different processes of spherulite formation in Aeolian Islands obsidian lavas.

GEOLOGICAL BACKGROUND

Recent activity in the central sector of the active Aeolian volcanic province has centred on rhyolitic activity, involving both effusive and explosive phases. On the islands of Lipari and Vulcano, rhyolitic eruptions (often accompanied by mafic enclaves, e.g. De Astis et al., 1997,

2013; Davì et al., 2009, 2010; Piochi et al., 2009; Forni et al., 2013, 2015) have occurred from ~43 ka (Tranne et al., 2002; Giocada et al., 2003; Lucchi et al., 2010, 2013; Forni et al., 2013), with the youngest eruptions occurring in historical time (Lipari - AD 776-1230; Keller, 2002; Forni et al., 2013; Vulcano - AD 1888-1890, Mercalli and Silvestri, 1891; De Astis et al., 2013). The oldest rhyolitic lavas are in southern Lipari (Punta del Perciato, Falcone, Monte Giardina and Punta di Costa lava domes; Figure 2), with younger domes in central and northern Lipari (Castello and Capo Rosso lava domes). At 8.7 ± 1.5 ka ago, activity shifted further north on Lipari, with the Pomiciazzo obsidian lava flow extruded from the Vallone del Gabellotto eruptive vent (Bigazzi and Bonadonna, 1973; Lucchi et al., 2010; Forni et al., 2013; Figure 2). The extrusion of the Forgia Vecchia and Rocche Rosse obsidian lava flows mark the end of this most recent period of activity on Lipari, with the Rocche Rosse obsidian flow dated by archaeomagnetic methods at $1,220 \pm 30$ AD (Tanguy et al., 2003), later revised to $1,230 \pm 40$ AD (Arrighi et al., 2006; Figure 2). Rhyolitic lava flows and domes have also formed on the neighbouring island of Vulcano (De Astis et al., 1997, 2013; Piochi et al., 2009). The Grotta dei Palizzi obsidian lava flow on the southern flank of the active Fossa cone (Vulcano) has been dated at 2.1 ± 0.3 ka (Voltaggio et al., 1995; De Astis et al., 2013; Figure 2), and, on the northern Fossa flank, the Pietre Cotte obsidian lava flow marks the latest outpourings of the 1739 AD activity (Frazzetta et al., 1984; Figure 2).

SAMPLING AND METHODS

Targeted samples span rhyolitic lava domes ranging in age from 43.0 ± 0.3 ka (Punta del Perciato, Falcone, Monte Giardina, Punta di Costa, Castello, and Capo Rosso lava domes on Lipari; Crisci et al., 1991; Lucchi et al., 2010; Forni et al., 2013), and younger lava flows (Pomiciazzo, Grotta dei Palizzi, Forgia Vecchia, and Rocche Rosse lava flows on Lipari and Vulcano; De Astis et al., 2010; Forni et al., 2013), to the 1739 AD activity (Pietre Cotte lava flow on Vulcano; Frazzetta et al., 1984) (Figure 2). Samples from younger obsidian lavas generally host better-preserved spherulites, and were therefore more extensively sampled. Samples were cut for thin section, and scanned for spherulite CSD analyses. A thickness of one device pixel in width was used. Processing programs *ImageJ* (Schneider et al., 2012), *CSDSlice* (Morgan and Jerram, 2006), and *MATLAB*® 6.1 commercial software were used to analyse all particles in a given image, including total spherulite count, spherulite phase abundance, and best fit spherulite shape. *CSDSlice* calculates a best fit for the aspect ratios, giving X, Y and Z values (defining the plane in a 3D space). *CSDCorrections 1.37* (Higgins, 2000, 2002, 2006), was used in order to calculate 3D crystal size distributions from 2D data.

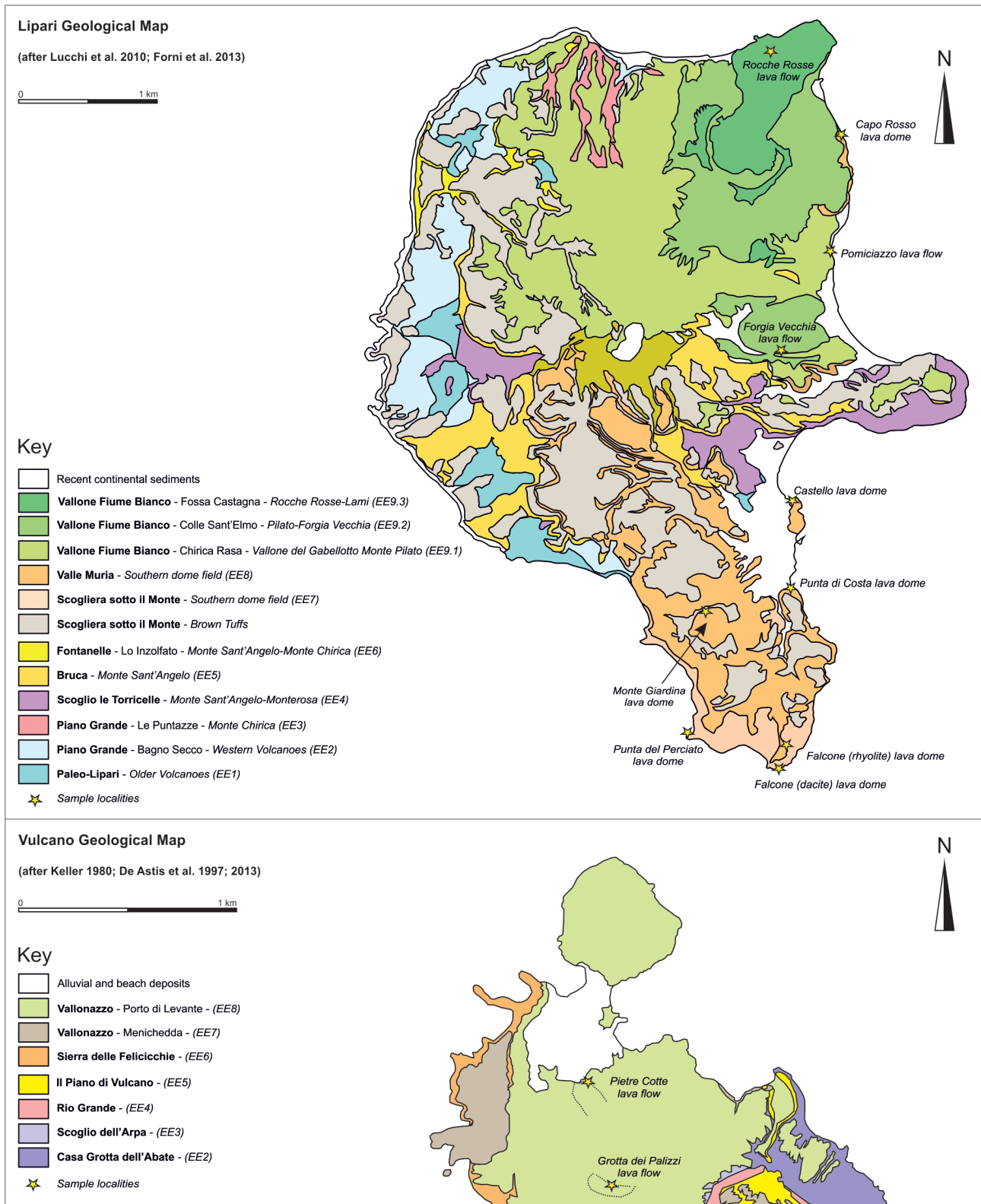


Figure 2. Geological maps and sample locations of obsidian lavas on Lipari and Vulcano (after Keller, 1970; Pichler, 1976; Lucchi et al., 2010; De Astis et al., 2013; Forni et al., 2013). Units classified by synthems, subsynthems and eruptive history as defined in the geological map of Forni et al. (2013), with synthems shown in bold text, subsynthems in normal text, and eruptive epochs in italics.

Glass and mineral major element analysis was undertaken at The Open University (UK) using a CAMECA SX 100 electron microprobe (EMP). The peak count per element for analyses was 15-30 s using a 10-20 μm defocused beam diameter for glass (a 1 μm wide beam was used for minerals at a count time of 40 s per element), with an acceleration voltage of 15 kV and a beam current of 20 nA for glass and minerals. Volcanic glass standards (VG-568 and KN-18) were routinely analysed as secondary standards. For XRD determination of individual compositional phases, spherulites were crushed and sieved, hand-picked using a binocular microscope, and powdered. The coarsely-powdered samples were placed on a flat disk sample holder, gently compressed, and scanned on a Bruker D8 Advance X-ray diffractometer, equipped with a VÅNTEC-500 detector at Keele University (UK), using $\text{CuK}\alpha$ radiation, a scan range from 5 to 90° (2 Theta), a 0.03° (2 Theta) step size and a data collection time of 10 s per step. Samples were prepared for Raman analysis by being cut to wafer-thin sections, $\sim 20 \mu\text{m}$ in thickness, and polished on both sides. Specific positions in the glassy groundmass, spherulites and brown rims were targeted using a high magnification petrological microscope. Spots of glassy groundmass free of microlites and alteration were targeted. Laser Raman spectroscopy analyses were performed with a Leica microscope coupled to a Renishaw Raman RM1000 system at the University of Southampton (UK), using a Renishaw NIR 780TF semiconductor diode laser (wavelength 780 nm) of power 25 mW. The spectrometer was set up in line with Renishaw recommendations for confocal operation of the spectrometer, calibrated using silicon with a known peak of 520 cm^{-1} . Spectra were obtained using an extended scan between 3700 cm^{-1} and 100 cm^{-1} , built up of 9 accumulative scans of 10 s, with a spot size of 1-2 μm , and processed using *SigmaPlot 10*.

RESULTS

Important characteristics in sampled obsidian lavas include variable deformational structures, degree of devitrification, spherulite morphology, and glass and spherulite geochemistry. These characteristics may impact (or trigger) spherulite nucleation, growth, and modification. CSD methods were utilised in order to assess physical and spatial features, while EMP, XRD, and Raman spectroscopy methods provide a means of geochemical characterisation of spherulites and host glass.

Textural observations

A number of well exposed lava flows and domes show brittle and ductile deformational fabrics, with multi-scale folding, stretching lineations, and sheared spherulites evident (Figure 3). Such deformational structures often coincide with areas of high spherulite density, and elongate

spherulites. Deformational fabrics are widespread across the Rocche Rosse lava flow, whereas deformed fabrics are more restricted to flow frontal regions and flow margins on the Pietre Cotte lava flow (Bullock, 2015). Structures in older lava domes are more difficult to distinguish due to weathering, glass alteration, and later pyroclastic coverage.

Optical observations of sampled obsidian lavas suggest distinct textural features within the obsidian samples, including widespread spherulites, a glassy groundmass containing microlites, a brown rim and clear halo surrounding most spherulites, and a crystalline mesh-texture found within some spherulites (Figure 4). Sampled obsidian across the data set is typically glassy black to microcrystalline grey in colour, and variably spherulitic (spherulites predominantly less than 1 mm in diameter, and often defining the planar foliation). Spherulites occur individually, interjoined, and in folds and fractures (Figure 4). Elongate spherulites can be observed in the Pietre Cotte, Rocche Rosse and Forgia Vecchia lava flows. These elongate spherulites tend to occur within or in close proximity of bands or folded outcrop. Generally, across all flows and domes, large (1-3 mm), spherical spherulites occur individually in zones of low shear or in zones of no banding. Spherulites in the Pomiciazzo and older sampled lavas tend to be much larger than those in younger flows, at 2-3 mm in diameter. The Punta del Perciato lava dome is highly devitrified, and individual spherulites are rarely evident to the naked eye.

Spherulite CSDs and growth periods

Clay et al. (2013) show that CSD quantitative textural analysis can be successfully applied to spherulitic samples, as the rate of isothermal spherulitic crystallisation was proven to be log linear in a given growth stage, despite being a non-equilibrium crystallisation process. The use of CSD methods is therefore an effective tool for the interpretation of spherulitic textures. Twelve CSD plots of spherulites from all lava flows and domes have been produced, with spherulites ranging in size from 0.1 mm (in younger flows), up to 8 mm (in older domes) (Figure 5). Overall, the sample set shows a kinked regression profile, with steeper, left-hand truncation at lower spherulite sizes, and a shallowing, more uniform regression to larger spherulite populations. The majority of plots show a higher degree of heterogeneity at smaller spherulite sizes, while some plots show little or no deviation from a log-linear profile. Older sampled domes exhibit larger spherulite maximum sizes and wider size ranges. Spherulites below 1 mm in size show a greater degree of complexity, evident by highly non-linear slope regressions. With the exception of samples from Castello, Punta di Costa and Monte Giardina, CSD slopes for sampled flows and domes often deviate from a straight line regression. Slopes show a distinct smaller population

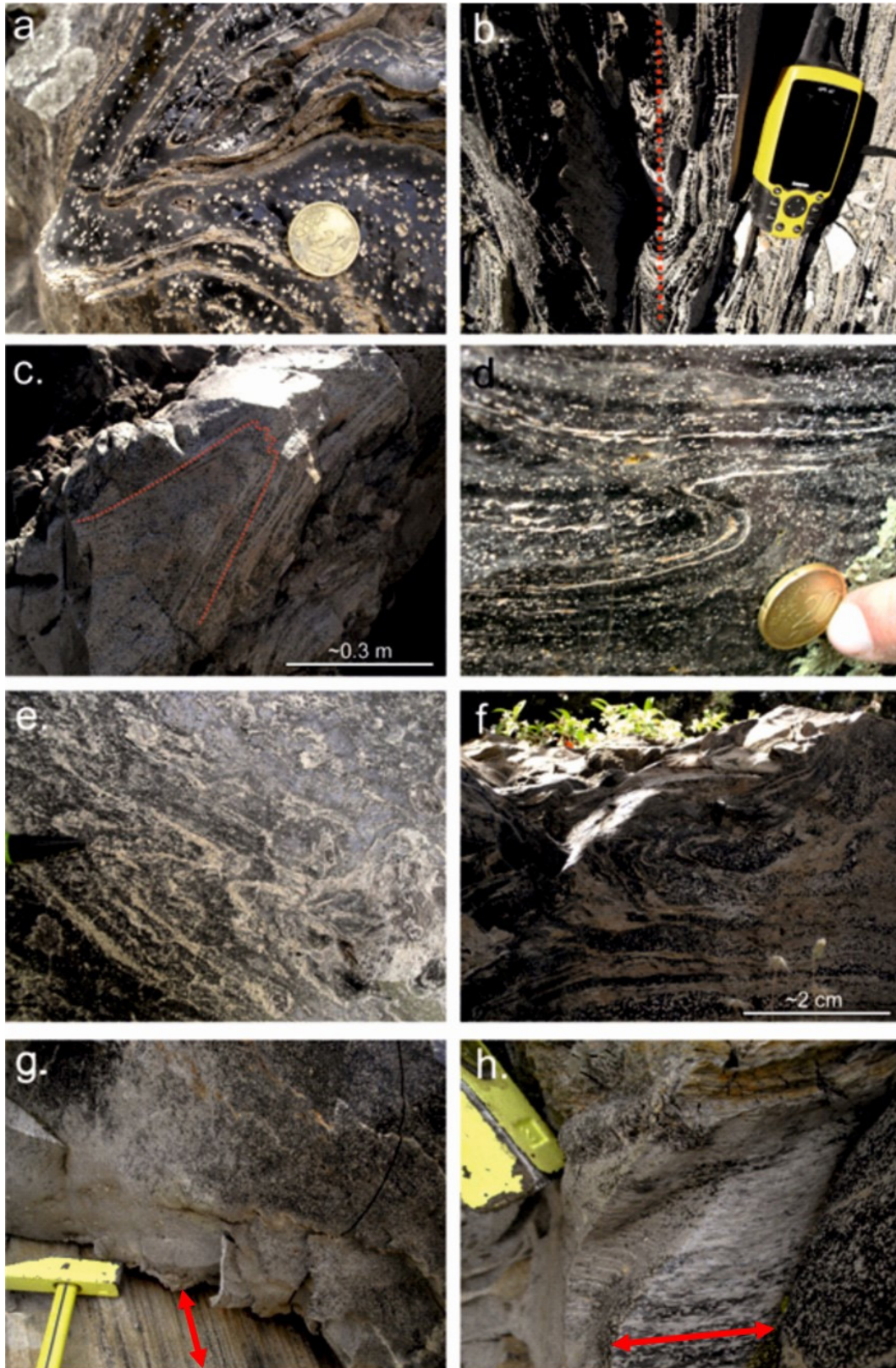


Figure 3. Evidence for spherulitic obsidian coinciding with areas of flow deformation, including (a)-(f) compressive folding (red line in b. highlighting the fold axial plane, and in c. representing fold outline) and (g)-(h) structures resulting from constrictional forces (arrows indicate stretching lineations).

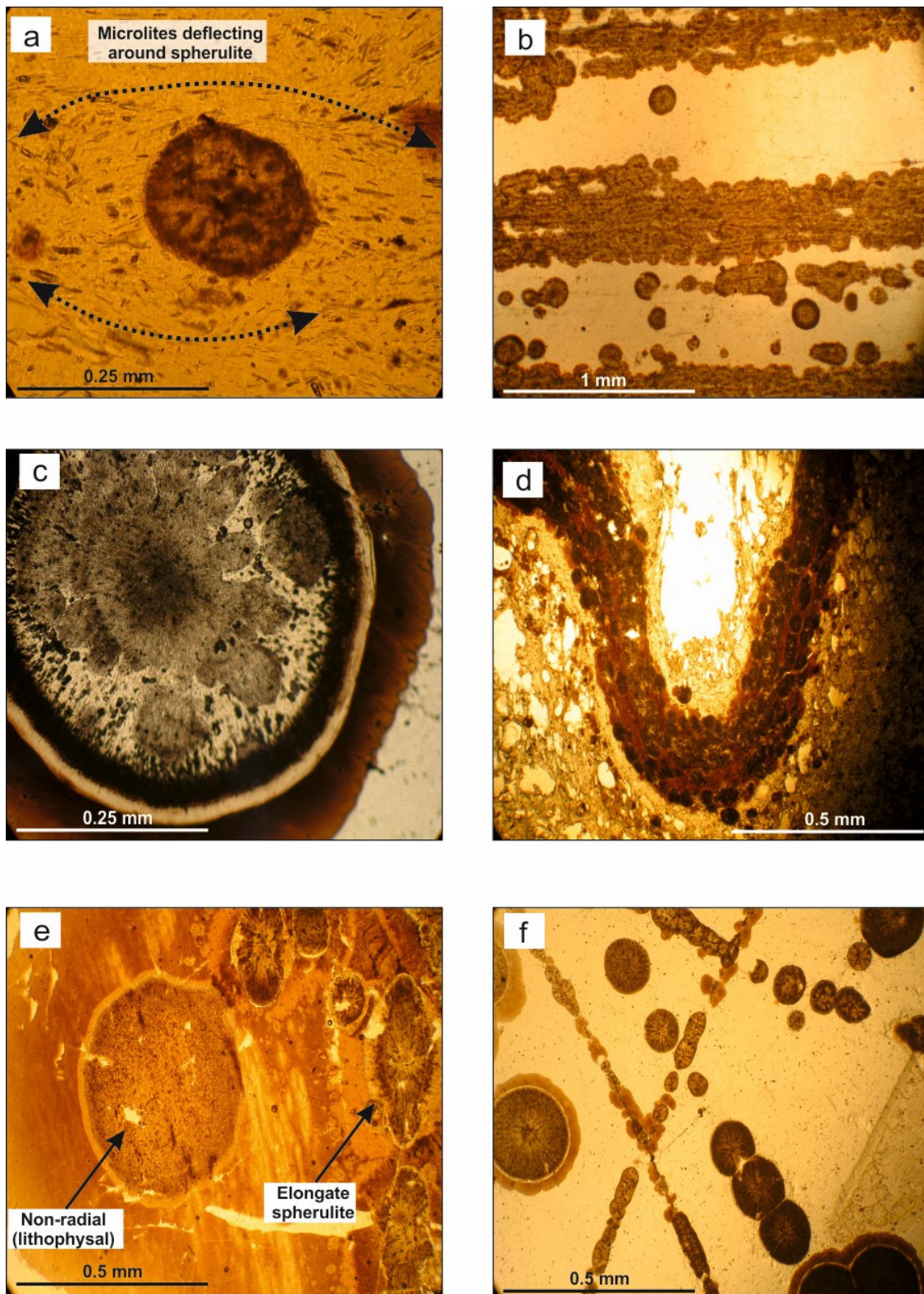


Figure 4. (a) Small spherulite (Pietre Cotte), with aligned microlites deflected. (b) Spherulites defining flow foliation, heavily concentrated within bands, with few spherulites in zones of low shear (Rocche Rosse). (c) Spherulite with radial centre, mesh-texture near edge, outer clear halo and surrounding brown rim (Falcone rhyolite). (d) Microscopic folding, with folded spherulitic-rich band (Rocche Rosse). (e) Microcrystalline spherulite (non-radial) and ellipsoidal, radial spherulites (Pomiciazzo). (f) Spherulites concentrated within conjugate fracture (Rocche Rosse).

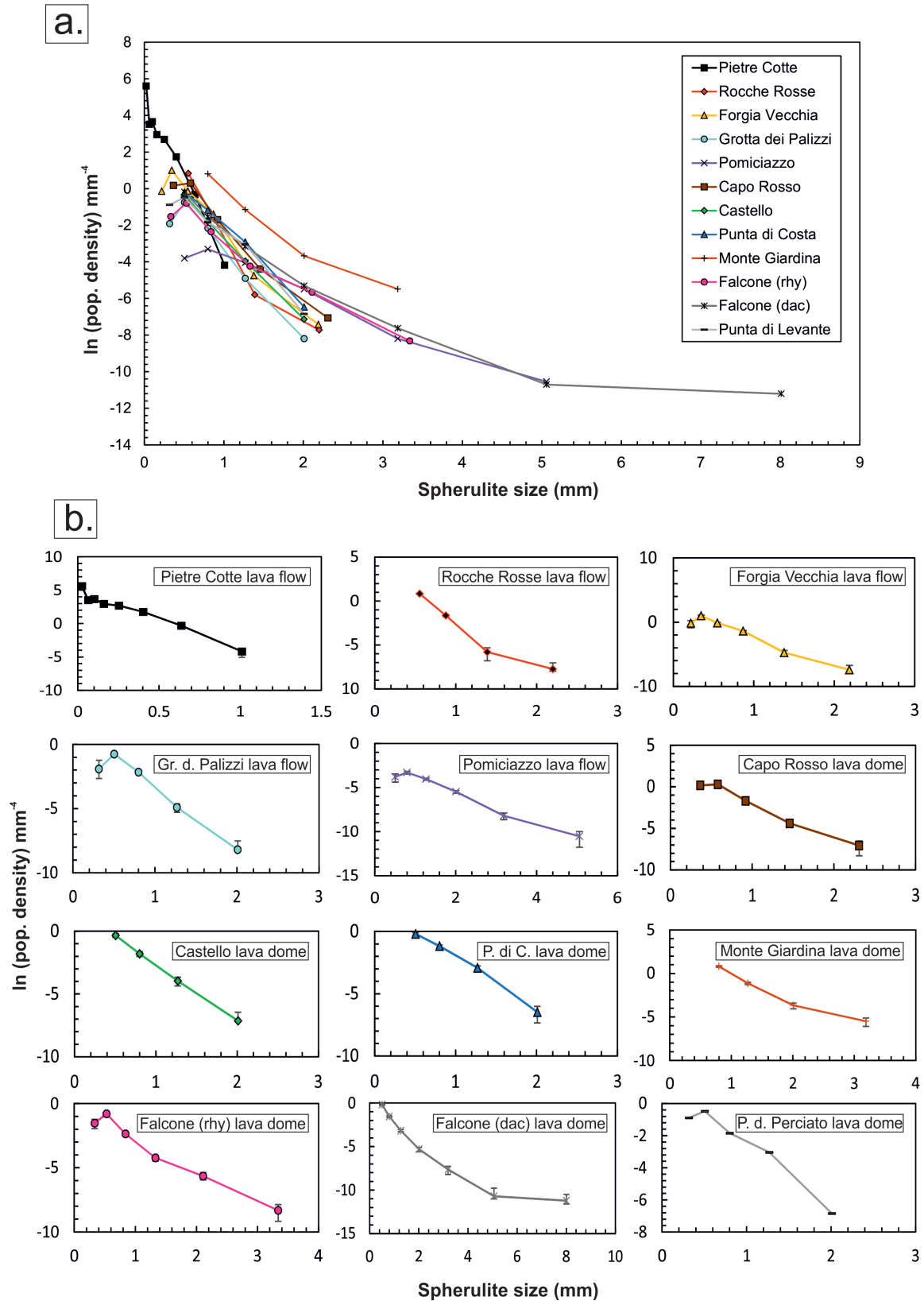


Table 1. Spherulite crystal size distribution (CSD) parameters for sampled lava flows and domes. Calculated growth periods for spherulite growth are also shown in the table (small and large size fractions removed and calculated through the main body of the CSD). Growth period based on a growth rate of $6.33 \times 10^{-7} \text{ m s}^{-1}$ (calculated by Castro et al., 2008).

Flow/dome	No. of spherulites	Y intercept	Slope value	R ² value	Short axis	Int. axis	Long axis	Sphericity (%)	Growth period (days)
Pietre Cotte	999	4.7997	-8.5831	0.9826	1	1.4	2.8	56.3	2.1
Rocche Rosse	451	5.2683	-7.9464	0.9998	1	1.05	1.15	92.6	2.3
Grotta dei Palizzi	154	1.7185	-4.9954	0.9966	1	1.05	1.05	98.4	3.6
Forgia Vecchia	395	2.4416	-4.6717	0.9838	1	1.15	1.25	90.3	4
Pomiciazzo	255	-1.9983	-1.7524	0.9859	1	1.05	1.05	98.4	10.4
Capo Rosso	446	2.3387	-4.2148	0.9793	1	1	1.15	91.1	4.3
Castello	130	1.8541	-4.5033	0.9992	1	1	1	100	4.1
Punta di Costa	431	2.1136	-4.1954	0.9933	1	1.05	1.05	98.4	4.4
Monte Giardina	326	2.3264	-2.5891	0.9489	1	1	1	100	7.1
Falcone (rhyolite)	576	-1.4589	-2.0418	0.9943	1	1.05	1.1	95.4	8.9
Falcone (dacite)	1086	1.2866	-3.3478	0.9912	1	1.05	1.05	98.4	5.5
Punta del Perciato	156	1.0329	-3.2901	0.9717	1	1	1	100	5.6

at <1 mm (Pietre Cotte, Forgia Vecchia, Grotta dei Palizzi, Pomiciazzo, Capo Rosso, Falcone rhyolite, and Punta del Perciato), a slope at intermediate sizes of 1-3 mm (Rocche Rosse, Falcone rhyolite, Falcone dacite, Punta del Perciato), and a shallower slope to larger spherulite sizes (Pomiciazzo, Falcone rhyolite, Falcone dacite).

Spherulite growth period was calculated using the CSD slope value (determined by *Microsoft Excel*), and assuming a constant spherulite growth rate of $6.33 \times 10^{-7} \text{ m s}^{-1}$ (calculated by Castro et al., 2008, in obsidian using water concentration profiles). The equation for calculating growth period is:

$$\text{Growth period} = -1 \div (\text{slope value} \times \text{growth rate})$$

(Marsh, 1988, 1998)

The average spherulite growth period across all individual slopes is ~7 days. However, larger spherulites may have formed due to modification/re-crystallisation processes occurring post-emplacement, and therefore may not be a true reflection of primary spherulite growth. Calculating growth periods from smaller and larger size fractions is also difficult as these are typically only defined by two bin sizes. Therefore, small and large size fractions

were excluded from calculations, and the slope value was extracted from the main linear body of the CSD profile, defined by three or more bin sizes. If slopes representing smaller and larger spherulites are excluded, the average growth period is ~5 days. Growth periods range across samples from 2 to 11 days (Table 1). A growth period has been calculated for smaller spherulites, though this is only tentatively suggested as the size fraction is defined by only two bins. Based on this estimation, smaller spherulites (distinct population < 1 mm) formed on average in less than a day. Intermediate spherulites (~1-3 mm) formed over an average of ~5 days (growth periods ranging 2-7 days based on the main body of the CSD profile). Linear slopes from Castello, Punta di Costa and Monte Giardina samples (spherulites typically ~1-3 mm in size) indicate that these spherulites also typically formed over ~5 days, with a range of 4-7 days.

Glass and spherulite geochemistry

All glass and mineral compositional data is provided in the Supplementary Appendix. The glass compositions range from rare dacitic/trachytic to typically rhyolitic. The minimum SiO₂ content is ~68 wt% (the Falcone lava dome and Punta del Perciato lava dome), and the maximum SiO₂

content is ~ 79 wt% (also from the Falcone lava dome) (Figure 6). The average SiO_2 content for the obsidian glasses is ~ 75 wt%. Al_2O_3 content varies from ~ 11 wt% to ~ 19 wt% (averaging ~ 13 wt%). Typical Na_2O content is ~ 3 - 5 wt%, with slightly higher Na_2O content in the Grotta dei Palizzi lava flow (~ 5 - 6 wt%). Brown rims surrounding spherulites show variable enrichment and depletion of SiO_2 , Na_2O and K_2O . Generally, spherulites show a high Si content, and a relatively high Al and K content, in specific phases (with a skeletal crystal arrangement). Electron microprobe spot analyses of spherulites reveal a mixture of phases, including glass (with silica values between ~ 70 wt% and ~ 78 wt%), and cristobalite (silica content of ~ 99 wt%), and orthoclase (silica content 60-67 wt%). Spherulites in all sampled obsidian flows and domes contain orthoclase and a silica phase. Plagioclase identified by EMP is primarily albitic.

Examination of Raman spectra for the glassy obsidian groundmass, spherulites (including the microcrystalline mesh-texture) and brown rims show minimal variation in notable peaks (Figure 7). Spectral results for the haloes show some slight variations. Sharp peaks represent the presence of rare crystalline material, and the broad humped regression is characteristic of amorphous material. Peaks at ~ 460 cm^{-1} and ~ 510 cm^{-1} confirm the presence of silica and orthoclase in spherulites and the glassy groundmass (microlites). There are also notable peaks at ~ 230 cm^{-1} and 300 cm^{-1} , with a relatively flat area between 800 cm^{-1} and 1200 cm^{-1} . Spherulites, mesh-textured interiors and surrounding brown zones show a similar trend to the glassy matrix, confirming the presence of rhyolitic glass in spherulites. Despite the differing optical properties (non-isotropic under crossed

polarised light), the brown rims show the characteristic profiles akin to a glass response. The main difference between glass and brown rim results is that although the brown rim shows an amorphous hump appearance, peaks tend to be sharper (more akin to crystalline material). The two major XRD peaks at 21.6° and 27.3° (2θ) correspond to α -cristobalite and orthoclase feldspar respectively, with minor peaks representing diopside (Figure 8). There are also peaks suggesting trace amounts of β -cristobalite, diopside and titanomagnetite, with peaks varying in occurrence and intensity across sampled lavas.

DISCUSSION

Textural variations in spherulites

Results suggest that the sampled obsidian suite across Lipari and Vulcano host spherulite populations of heterogeneous textural characteristics. This is evident in both petrographic observations and kinked CSD profiles (Figure 6). The upward kinking profile of the CSD plots indicates increasing heterogeneous spherulite nucleation, e.g. faster or denser spherulite nucleation at a later stage. Distinct CSD slopes reflect individual spherulite populations, nucleating and growing at different times and at different rates. This results in texturally heterogeneous spherulite populations within an obsidian lava. In some instances, there is evidence for a downturn at the smaller spherulite sizes (Figure 6). Petrographic observations suggest that spherulites do not occur below the resolution limit, and smaller spherulites are optically distinct compared to larger spherulites. For spherulites, the upward inflection suggests a clustering or coalescence of large spherulites, the result of localised

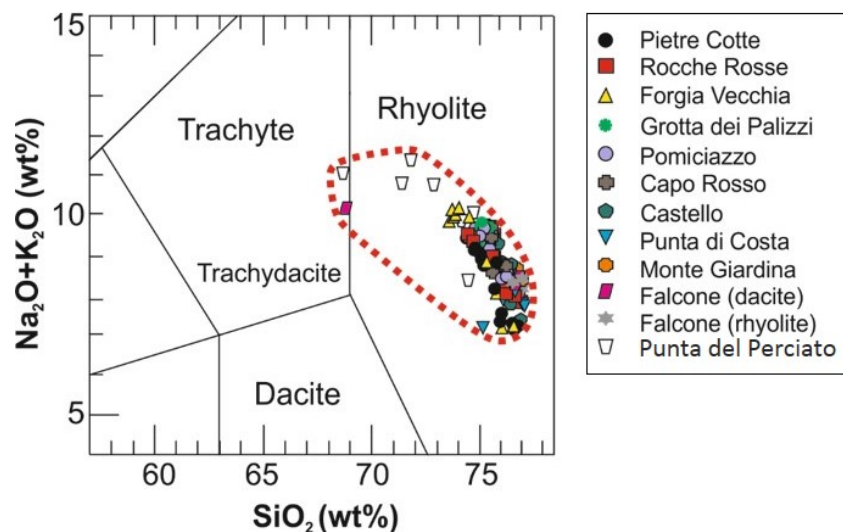


Figure 6. Total Alkali-Silica (TAS) diagram (adapted from Le Bas et al., 1986) of glass compositions across sampled lavas. The majority of points are clustered together at 73-76 wt% SiO_2 and 7-9 wt% $\text{Na}_2\text{O} + \text{K}_2\text{O}$ (shown by red field).

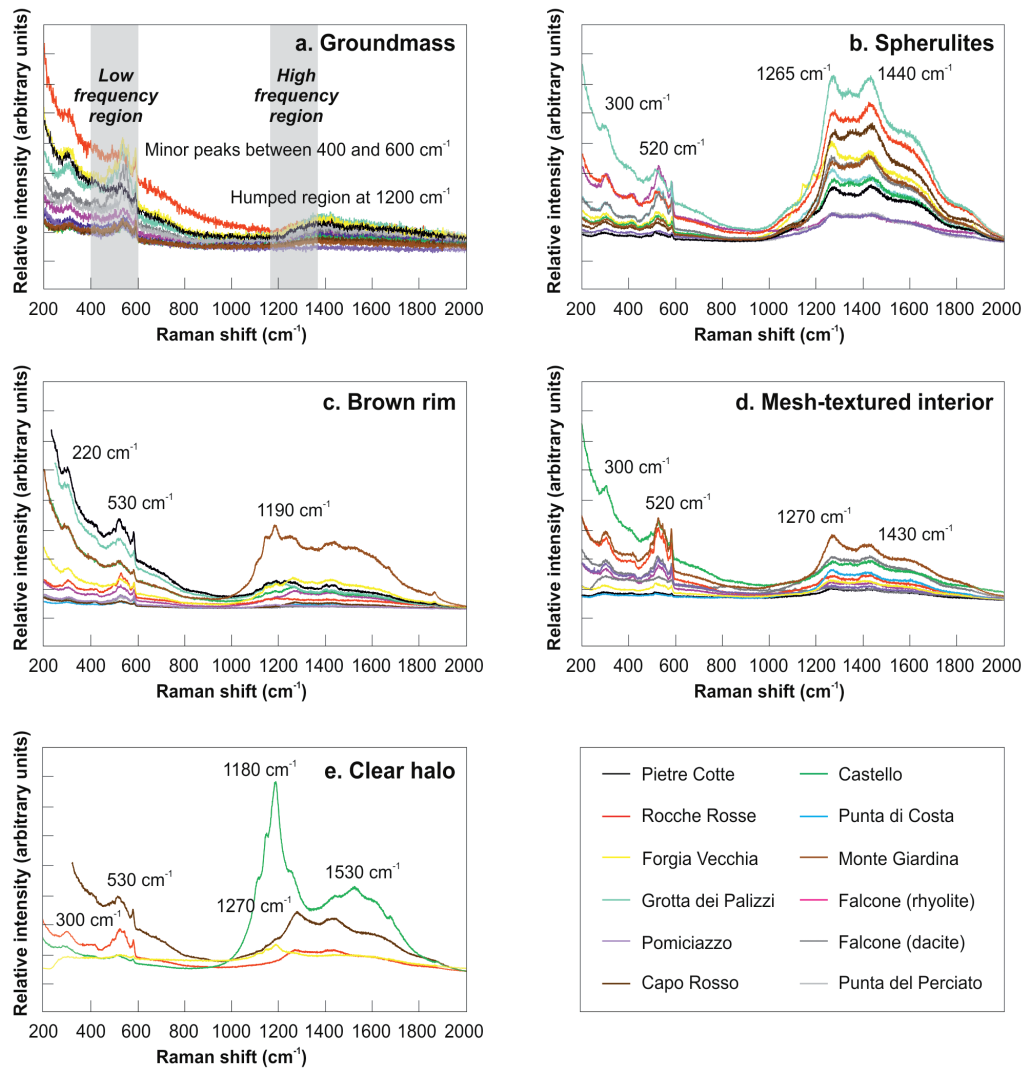


Figure 7. Raman spectrum and peaks for (a) obsidian glass groundmass, (b) spherulite interiors, (c) brown rims surrounding spherulites, (d) mesh-like textures found in some spherulites, and (e) clear haloes immediately surrounding some spherulites.

high nucleation rates and/or the product of prolonged devitrification. Later-stage vapour-phase crystallisation overprinted or re-crystallised earlier formed spherulites, evident by lithophysal textures with an original spherulite in the centre, and resulting in a greater population density of larger spherulites. Samples exhibiting a simple log-linear regression slope indicates a single nucleation event, likely to be the result of constant undercooling. Generally, younger flows have shorter growth periods, and older domes have longer growth periods. This may be a reflection of the size of spherulites being a function of a constant growth rate. However, spherulites are likely to grow at a different rate in a melt compared to those growing in a semi-solid or solid obsidian state. Spherulites may also have grown at different rates from different spherulite-forming processes, i.e. spherulites

growing at high temperatures as a response to enhanced undercooling, and spherulites nucleating and growing as a result of deformation. Calculated growth periods for older samples may also take into account the effects of late stage infilling phases (Holzhey, 2001). Notable exceptions to spherulite size being primarily a function of time are large spherulites in samples from the Pomiciazzo lava flow and the Falcone dacite lava dome. As well as time, spherulite growth is primarily a function of temperature and flow rheology. At temperatures approaching melting point, molecular stems settle for longer on the growing crystal face, thus allowing spherulites to grow faster (Lauritzen and Hoffman, 1973; Hoffman et al., 1975; Hoffman and Miller, 1997). Growth may also relate to viscosity, with high T_g viscosities inhibiting diffusion, and thus reducing spherulite growth rate. This suggests that larger spherulites

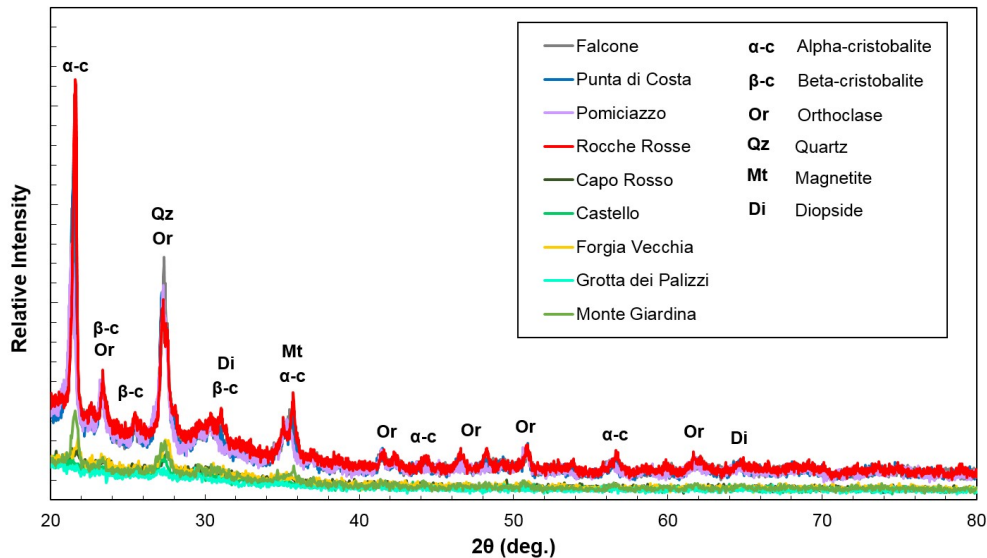


Figure 8. X-ray diffraction (XRD) results and peaks for spherulites from sampled lavas. Notable peaks include several α -cristobalite and orthoclase peaks (main constituents), with minor magnetite, β -cristobalite, and diopside peaks.

observed in Pomiciazzo and Falcone dacite samples may be a result of differential systems exhibiting prolonged period of higher temperatures and lower viscosities.

The regular and highly recognisable occurrence of densely spherulitic glass coinciding with flow deformation, and the uniform occurrence of elongate and/or sheared spherulites, supports a link between flow deformation and spherulite nucleation. Flow-induced crystallisation, also referred to as flow-enhanced nucleation or shear-enhanced nucleation (Graham and Olmsted, 2009), is a well understood phenomenon in polymer sciences as a trigger for spherulite nucleation, but is thus far a neglected concept in a geological context. Elongation in an extensional flow regime has been cited as a strong stimulus for polymer crystallisation, causing molecules to orient and stretch in the direction of extension, facilitating the process of flow-induced crystallisation (Janeschitz-Kriegl et al., 2003; Kornfield et al., 2002; Stadlbauer et al., 2004). Shear flow can induce crystallisation, requiring very large generation of strain (Derakhshandeh and Hatzikiriakos, 2012). Extension and shear flow-related crystallisation may therefore be applicable in highly deformed obsidian lavas. Nucleation depends on the change of the free energy difference between the crystal and the melt phases; the degree of order in the melt phase increases under shear conditions, and the free energy difference increases. The entropic penalty for crystallisation is lowered and, therefore, the nucleation rate increases (Keller and Kolnaar, 1997; Coccorullo et al., 2008). This leads to a high number density of spherulites as a result of deformation, which is evident in field observations in obsidian lavas. This

may suggest that, in obsidian lavas, spherulite nucleation may increase with flow deformation (during flow), and remain constant in areas of low (or no) strain (e.g. in areas unaffected by strain during flow, or following flow cessation). This results in high spherulite populations in areas of high strain (and spherulites may be sheared or elongate as a result, also observed in the Rocche Rosse lava flow by Clay et al., 2013), and lower populations of (undeformed) spherulites in zones of low strain, allowed to grow and form more typical spherical shapes.

Elongate spherulites (elliptical as opposed to bowtie shaped) spatially found within close proximity to deformational structures in sampled flows and domes may have formed while lava was behaving in a ductile manner (within T_g). This is similar to how vesicles and enclaves preserve lava flow strain, based on their shape and size alteration from spherical to elongate (Polacci and Papale, 1997; Rust et al., 2003; Ventura, 2004; Iezzi and Ventura, 2005). The onset of this spherulite nucleation and growth may therefore have been triggered (or enhanced) by flow-induced crystallisation (i.e. as a result of deformation) at high T_g temperatures. This is supported by polymer experiments whereby flow-induced crystallisation occurs near melting point (above T_g or high T_g transition) (Binsbergen, 1966). The stress created by both spherulite nucleation and that acting due to lava flow emplacement modifies the spherulite shape from spherical to elongate. In zones of high strain, such as at flow margins or where individual flow lobes overlap, spherulites are completely sheared into fully spherulitic bands, or microlite-rich flow bands (which may have resulted from flow-induced crystallisation) became spherulite-rich due to microlites

acting as spherulite nuclei. There is also clear evidence for spherulites nucleating and growing within brittle fractures, such as conjugate fractures and tension gashes. This suggests that spherulites have preferentially nucleated in these fractures while the flow is behaving in a brittle manner, i.e. at low or below T_g temperatures in localised regions of flow or due to high strain rate, as a response to deformation.

Geochemical variations in spherulites

The presence of trace amounts of high temperature β -cristobalite suggests that spherulites were forming at least higher than the β - α temperature transition (~ 270 - 200°C ; Wright and Leadbetter, 1975; Downs and Palmer, 1994; Swainson and Dove, 1994; Damby et al., 2014), as a quench phenomenon (Ewart, 1971). Spherulites were likely to begin forming early, with high temperature β -cristobalite converting to low temperature α -cristobalite (in some instances, incomplete conversion resulting in both α - and β - cristobalite). These phases are identified in the XRD peaks. Variations in the intensity and shift of peaks in the clear surrounding haloes results may suggest some degree of compositional variation (evident in EMP results), and may relate to different processes forming the haloes, e.g. secondary crystallisation infilling (Breitkreuz, 2013), or Fe-redox shift (Castro et al., 2009).

Results suggest some depletion and enrichment of major elements within spherulites relative to the glassy matrix, and a surrounding rim. Following initial spherulite growth, cavities at the interface of the spherulite and surrounding material may open out, which act as a nucleation surface for a second phase of crystallisation (Breitkreuz, 2013). Where the surrounding halo does not show a crystalline interior, the surrounding colourless halo may be enriched in OH groups and depleted in ferric iron, produced by a redox front (driven by magnetite crystallisation) that originated from the spherulite margin (Castro et al., 2009). Low-temperature components which are rejected at the crystal-melt interface, assisted by a low diffusion rate, form an impurity layer (Keith and Padden; 1963; Lofgren, 1971b). This impurity layer may be identifiable in Aeolian Islands obsidian samples as the surrounding colourless halo or brown rim evident across all lava bodies. The presence of glass within spherulites is reported here, and also noted elsewhere (e.g. Richnow, 1999; Ryabov and Grib, 2005; Castro et al., 2008, 2009; Seaman et al., 2009; Gardner et al., 2012). Seaman et al. (2009) conclude that the preservation of glass within spherulites highlights the critical influence of water concentration on the likelihood of quenching versus crystallisation during undercooling of a melt. Richnow (1999) also suggests that the presence of glass within spherulites is an indicator of incomplete crystallisation and spherulite formation above T_g temperatures.

Processes of spherulite formation in Aeolian Islands obsidian lavas

As rhyolitic lava cooled, it passed from liquid-like behaviour to solid (glass)-like behaviour (Gottsmann and Dingwell, 2001b). Nucleation and spherulite formation originally occurred as a response to undercooling, and continued as a down-temperature growth continuum (Clay et al., 2013). In the majority of sampled lavas, microlites grew first and were later overgrown by spherulites. Therefore spherulites mainly formed at lower temperatures than microlite crystallisation. However, smaller spherulites sometimes show microlites deflected around them, indicating that spherulites grew at higher microlite crystallisation temperatures. Spherulites which exhibit a radial centre and non-radial outer zone preserve evidence for both primary spherulite crystallisation and lithophysae forming processes. Spherulites are often surrounded by a brown rim, but there are smaller spherulites which do not show this rim. These spherulite rims may have formed within or below T_g as a solid state reaction (Castro et al., 2009).

Textural observations, geochemistry and CSD plots suggest that multiple spherulite populations exist across obsidian flows and domes on Lipari and Vulcano. Each population can be classified according to their textural heterogeneity (Figure 9). From the digitised spherulites, a representative shape for each spherulite type was created using the CSDSlice program of Morgan and Jerram (2006). X, Y and Z values were plotted in MATLAB[®] 6.1 in order to provide a representative 3D sphere for each spherulite type (Figure 9). Here, the X, Y and Z data were scaled with the radius, with the centroid of sphere representing 0 on each of the three axes. The equal square axes allow for direct comparison of representative spherulite type shapes, and show true elongation. Representative shapes show that early, high temperature spherulites are more deformed than later, lower temperature spherulites, peaking at high T_g temperatures, with deformation waning towards lower T_g temperatures and below. Initial spherulite nucleation and growth may have been triggered by enhanced undercooling above T_g . This results in spherulites with radial interiors. These can be classified as type 1 spherulites. Microlites were deflected by these high temperature spherulites. Type 1 spherulites are typically the smallest spherulite types (responsible for left hand truncation in spherulite CSDs). The formation of spherulites during emplacement at high T_g transition results in elongate (or sheared) spherulites during ductile flow deformation, which are classified here as type 2 spherulites. The occurrence of isolated, spherical spherulites (larger than type 1 spherulites) with a brown rim suggests that some spherulites formed in areas of low strain. These are classified as type 3 spherulites. Other spherulites grew at temperatures within and below T_g , and concentrated within fractures or bands (deformational

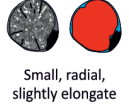
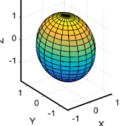
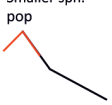

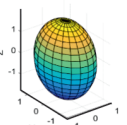

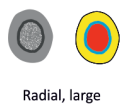
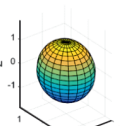

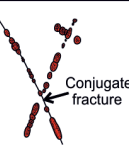
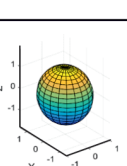
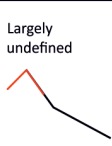
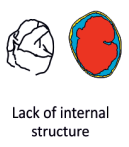
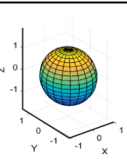

Type	Image	Representative shape	Typical size	Features	Composition	Growth period	% of total sph.	% of total sph. area	Typical CSD slope	Formation
1	 Small, radial, slightly elongate		< 1 mm	Radial interior, sub-spherical, occasional mesh-textured rim, microlites deflected	α - β cristobalite transition (some β crist. fixed in matrix), orthoclase, glass, diopside, titanomagnetite	Less than one day	19.3%	3%		Enhanced undercooling at high T temperatures above T_g (temperatures similar to microlite crystallisation)
2	 Elongate		1-2 mm	Elongate shape, vague radial interior, associated with deformational fabrics, brown rim, mesh-texture	α cristobalite, orthoclase, glass. Zones and pockets of pure crystalline silica	2-7 days	10.2%	13.3%		Constant degree of undercooling and deformational trigger. Nucleation during high T_g temperatures
3	 Radial, large and spherical		1-3 mm	Radial interior, sub-spherical, associated with zones of low shear, clear halo, mesh-textured rim, think brown rim, microlites unaffected	α cristobalite, orthoclase, glass. Zones and pockets of pure crystalline silica	2-7 days	51.9%	39.4%		Constant enhanced undercooling and deformational trigger (low intensity), mid to sub T_g temperatures
4	 Conjugate fracture		<1 mm	Concentrate within joints and fractures, often interjoined, no halo or brown rim, radial and no zonation	α cristobalite, orthoclase, glass	-	5.3%	2.8%		Deformation and preferential nucleation within constrained faults and fractures, below T_g , nucleation and restricted growth (spatially-controlled)
5	 Lack of internal structure		>1 mm	Partially radial or microcrystalline interior, non-continuous or poorly defined spherulite edge, mesh-like texture	α cristobalite, quartz, orthoclase, glass	-	13.4%	41.7%		Modification of previous types at vapour stage temperatures and below T_g , secondary alteration of spherulites and cavities. Otherwise classified as lithophysae

Figure 9. Spherulite type classification scheme based on textural characteristics, composition, and CSD plots. In the digitised images, red areas represent a primary spherulite phase, blue is a secondary infilling phase, and yellow is a reaction rim.

trigger), classified as type 4 spherulites. Hollow, partially-hollow, or spherulites containing tensional voids and cavities later became infilled or re-crystallised, resulting in internal spherulite colour zonation and pockets of crystal phases. These are classified as type 5 spherulites (modification of other spherulite types), and may only partially contain a radial texture or no radial texture at all. During ascent, eruption and emplacement, lava is likely to cross T_g many times, resulting in breaking, flowing, and healing in repetitive cycles (Tuffen et al., 2003; Tuffen and Dingwell, 2005; Vasseur et al., 2013). As a result, spherulites types formed concomitantly in different parts of the obsidian lava body, relating to temperature variations and different processes at work (e.g. at higher temperatures in the core of the obsidian lava, or as a result of flow-induced crystallisation).

Though spherulite formation is a function of processes such as undercooling and deformation, it is unlikely to be fully time constrained. Repeated fracturing and healing of rhyolite (Gonnermann and Manga, 2003; Tuffen et al., 2003; Tuffen and Dingwell, 2005; Vasseur et al., 2013),

and second boiling (separation of a gas from a liquid phase) also occur (Manley and Fink, 1987; Westrich et al., 1988; Sisson and Bacon, 1999; Tuffen et al., 2012), which may re-initiate crystallisation. Therefore, differing spherulite types may nucleate and form simultaneously (Figure 10). Rheological properties such as elasticity and temperature are intrinsically linked. Such rheological parameters will differ across lava flows and domes during and after emplacement. It can therefore be assumed that deformation is intrinsically linked with temperature and different spherulite types form at the same time. In an obsidian flow, for instance, it will be cooler at flow margins and hotter in the centre of the flow. Therefore, spherulites forming via high temperature undercooling in one part of obsidian lava (such as the core of the flow) do so concomitantly with spherulites forming during T_g in other parts of the lava (by flow-induced crystallisation at margins). Type 5 spherulites, characterised by modification and filling of cavities, may occur during and after vapour-activity and fracturing. Such recrystallisation textures developed due to primary cooling, and continued

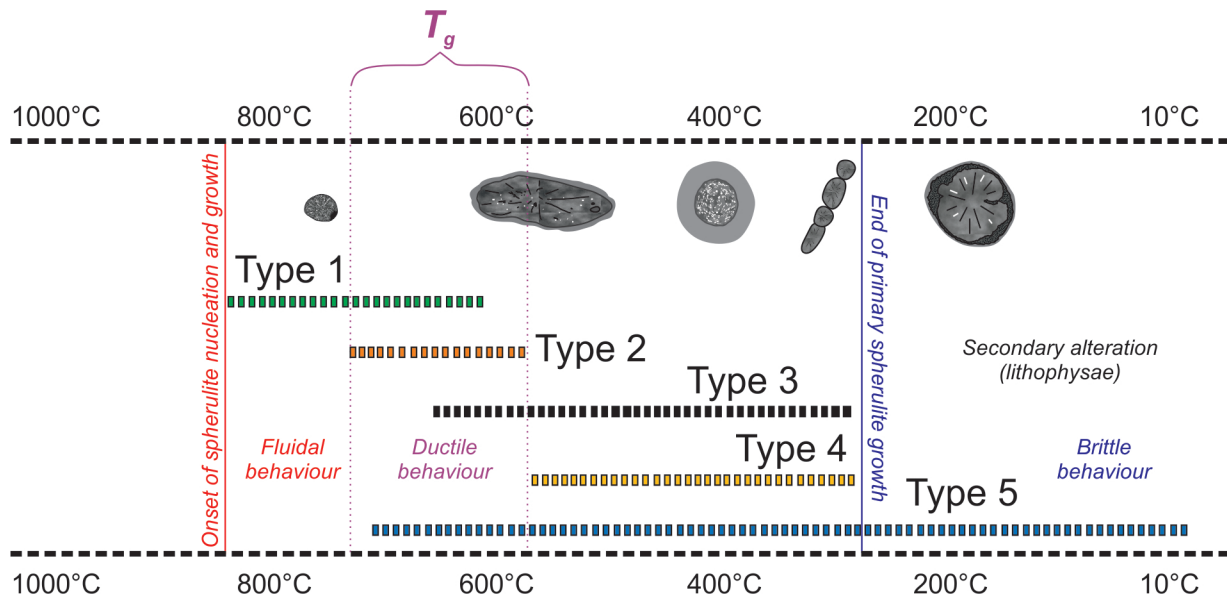


Figure 10. Temperature model for the formation of classified types of spherulites (texture process timeline modified from Breitreuz, 2001). Model demonstrates that spherulite type classification is rheologically controlled, and spherulite types may form simultaneously in different areas of lava. Indicated T_g range is based on previous estimations (Ryan and Sammis, 1981; Manley, 1992; Davis and McPhie, 1996; Watkins et al., 2008; Gardner et al., 2012; Clay et al., 2013).

throughout the cooling history of the lavas. Repeated fracturing and healing also means that type 4 spherulites cannot be fully time-constrained, but are restricted to areas of flow that are undergoing a brittle flow regime.

Spherulite growth and modification

A total of 5405 spherulites were digitised according to their designated type, from samples spanning all sampled lavas (Figures 9 and 11). Type 3 showed the highest population of spherulites (51.8%). Types 1, 2 and 5 show a similar percentage of the total population (19.3%, 10.1% and 13.4% respectively), and type 4 was the lowest percentage of the total population (5.3%). Despite type 3 spherulites having the highest population of spherulites across samples, type 5 spherulites covered the highest area fraction (43.3%). Types 1 and 4 covered the smallest percentage area fraction (both 3.3%). These results show that types 3 and 5 dominate spherulite populations in obsidian, both in terms of number of spherulites and area (and volume) covered. Smaller spherulites (typically type 1) exhibit the shortest growth period, estimated at less than a day at high T_g /above T_g temperatures. Spherulite types 2 and 3 are considered to have started growing at high T_g temperatures. The average growth period of ~4 days is comparable to that calculated for Rocche Rosse spherulites (also 4 days) at temperatures $\geq 800^\circ\text{C}$ by Clay et al. (2013). Type 5 (larger spherulites) have longer calculated growth periods, but processes of modification/re-crystallisation may be responsible for larger crystal sizes, independent of growth period.

Type 4 spherulites are often spatially restricted, and thus growth periods were not calculated. This is further substantiated by plots of individual spherulite populations and histograms in Figure 11. Figure 11c-h show that types 3-5 are skewed towards more spherical spherulites, while types 1 and 2 (in particular, type 2) show a more even (and elongate) distribution. Type 1 (high temperatures above T_g) spherulites show some elongation, while type 2 spherulites (high T_g temperatures) show the greatest degree of elongation. Spherulite types 3-5 show a more ideal spherical shape. These observations demonstrate some shape variation within each type, and that flow deformation was at its peak at high T_g temperatures, evident by highly deformed type 2 spherulites.

CONCLUSIONS

Examination of spherulites across the exceptionally-preserved obsidian lava domes and flows of Lipari and Vulcano show textural and geochemical heterogeneity, a result of different spherulite-forming processes. Such recrystallisation textures developed due to primary cooling, and continued throughout the cooling history of the lavas. Processes occur pre-, syn- and post-emplacement, across glass transition temperatures (T_g). This results in five categorised spherulite types:

Type 1 spherulites - small, radial and slightly deformed spherulites, formed due to high degrees of undercooling at high T_g transition temperatures or above.

Type 2 spherulites - elongate spherulites with a surrounding brown rim, formed due to deformation (tend

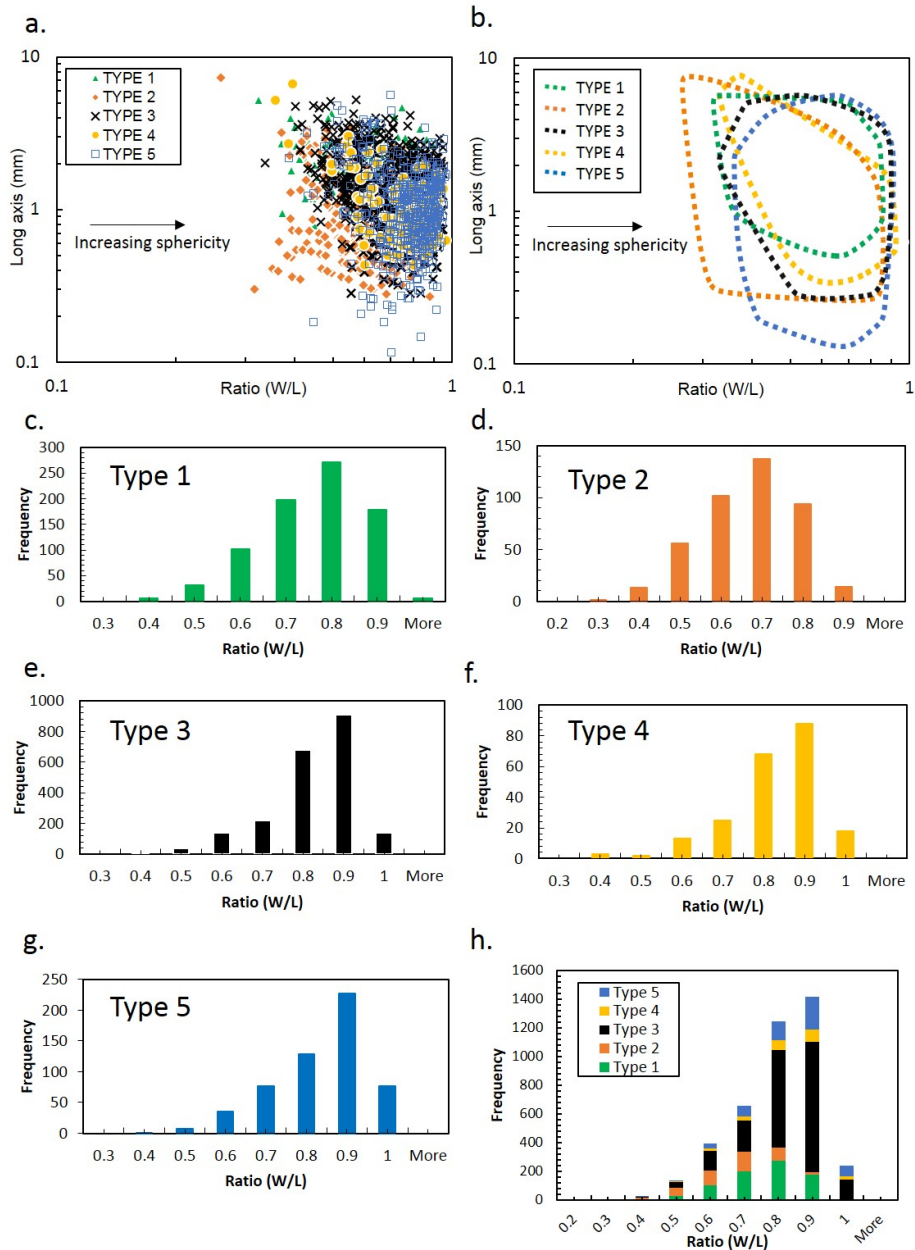


Figure 11. Individual spherulite parameters (categorised by types). (a) Plot of long axis vs. ratio of length and width (W/L) for all spherulites. (b) Representative Fields of long axis vs. W/L ratio, showing that the majority of type 2 spherulites are more elongate than other types. (c-g) Histograms for spherulite types and (h) stacked histogram of all types.

to form in shear zones) and/or enhanced undercooling.

Type 3 spherulites - large, spherical spherulites, often occurring individually in zones of low shear regions, formed by enhanced undercooling or deformation in low T_g transition or below T_g .

Type 4 spherulites - concentrated within fractures and faults, formed following solid state deformation.

Type 5 spherulites - partially- or fully-modified spherulites and lithophysae, with a non-radial interior, formed by vapour-infilling (low T_g transition or below).

Observations highlight the role of enhanced undercooling, deformation, flow-induced crystallisation, and post-emplacement modification in the development of spherulites across T_g temperatures, and considerable spherulite heterogeneity in obsidian lavas.

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