



## Extensional magmatism in a continental collision zone, Tafresh area, western central Iran: structural, geochemical and mineralogical considerations

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### ABSTRACT

The Tafresh area, located on the middle part of Urumieh-Dokhtar magmatic arc, is a suitable area for considering the effects of tectonic structures on development of the volcanic facies in the magmatic provinces. The main structure in the area is a dextral fault (so called Rahjerd Fault) that has produced the local extensional basins and also the parallel dyke swarms (or feeder dykes) on its both sides. So, the pyroclastic deposits, basaltic-andesitic lava flows, parallel dyke swarms were appeared by explosive subaqueous and then effusive subaerial eruptions and eventually a dioritic stock intruded in the volcanic pile. Petrologically, the magmatic rocks are belonged to the calcalkaline suite that had been changed and evolved in the magma chambers of local extensional basins in a collisional tectonic setting. The mineral chemistry and geochemical modeling as well as the coexistence of different mineral paragenesis of plagioclases ( $An_{47-72}$  and  $An_{17-37}$ ), pyroxenes (diopside and pigeonite) and amphiboles (magnesiohastingsite and magnesiohornblende) reveal that the AFC process is a dominant process in the magma chamber. Also, according to geothermobarometric calculations, the investigated volcanic rocks of the area could be grouped into two types: one type with higher P-T (>6Kbar and about 900 °C) including Eocene pigeonite-magnesiohastingsite bearing andesites (PHA) and Miocene andesites and other with lower P-T (1-4 Kbar and 600-800 °C) including the Eocene diopside-magnesiohornblende bearing andesites (DHA) and the dioritic stock.

Keywords: parallel dykes; extensional basins; AFC; geothermobarometry; Tafresh; central Iran.

### INTRODUCTION

Urumieh-Dokhtar magmatic arc (UDMA) located on the western Iran (Figure 1), was formed and evolved by collision of the Arabian and Iranian continental lithospheric plates following closures of Neotethys ocean during Alpine orogenies (e.g., Berberian and King, 1981; Alavi, 1994; Agard et al., 2005; Vincent et al., 2007; Horton et al., 2008; Morely et al., 2009). The NW-trending UDMA is the SW-border of central Iran microcontinent (Alavi, 1994) with a 1700 Km in length

from Anatolian Fault in Turkey to Oman Line and 100-150 Km in width. It is characterized by two prominent lithologic and structural characteristics:

1) It contains a thick sequence (up to 3-4 Km) of volcanoclastic deposits, basic-acidic lava flows and subvolcanic intrusions with a calcalkaline affinity. Although the oldest magmatic event in UDMA is attributed to Jurassic (Jazi et al., 2012) or early Cretaceous (Alavi, 1994) and continued up to Quaternary (e.g., Agard et al., 2005; Omrani et al., 2008), most researchers (e.g.,

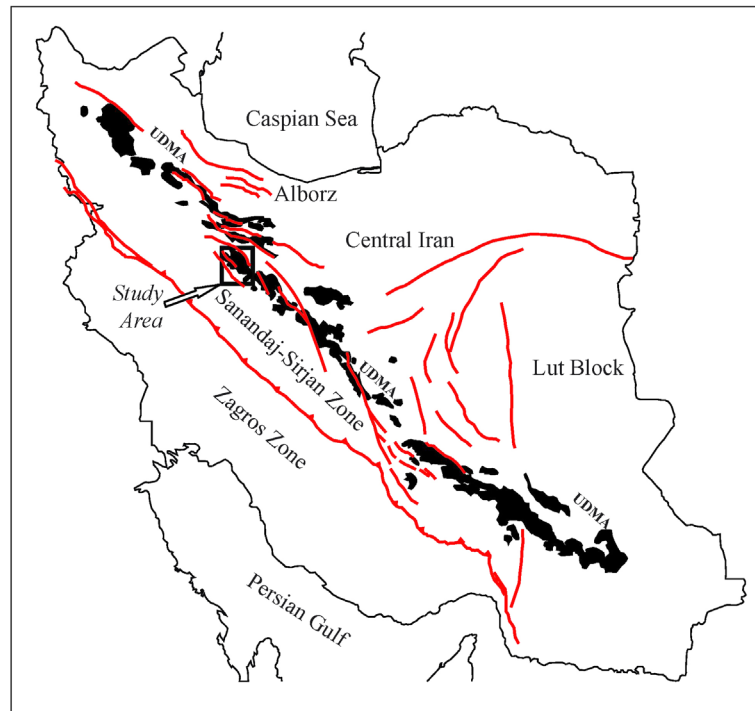


Figure 1. Situation of the study area (rectangle) in UDMA.

Stöcklin, 1974; Farhoudi, 1978; Omrani et al., 2008) suggest that the most widespread event in UDMA was occurred mostly in Eocene.

2) UDMA coincides to a NW-trending strike slip fault zone including North Tabriz Fault, Qom-Zefreh Fault and Dehshir-Baft Fault. It seems that this fault system played an important role in displacement of structures and likely volcanism in UDMA.

Thus, this study focuses mainly on the interplay between volcanism and tectonism in a magmatic arc via facies analysis, geochemical considerations and structural analysis.

#### LITHOSTRATIGRAPHY AND PETROGRAPHY

Urumeih-Dokhtar volcanic zone of Schröder (1944) is an Andean-type magmatic arc that has been active from late Jurassic to the present (e.g., Berberian and King, 1981; Berberian et al., 1982). However, Silurian volcanism has been reported in it recently (Tabatabaeimanesh et al., 2011).

The UDMA is mainly composed of voluminous calcalkaline and K-rich alkaline intrusive and extrusive rocks with associated pyroclastic and volcanoclastic successions (e.g., Jung et al. 1976; Ahmad and Posht Kuhi, 1993; Agard et al., 2005; Omrani et al., 2008; Dilek et al., 2010). The oldest rocks in the UDMA are calcalkaline intrusive bodies, which cut across late Jurassic formations and are overlain unconformably by lower Cretaceous

fossiliferous limestones. On the other hand, the youngest rocks in this zone are Plio-Quaternary lava flows and pyroclastic deposits (Berberian and Berberian, 1981).

Ballato et al. (2011) suggest a two-stage collision process for Arabia-Eurasia collision zone, involving the “soft” collision of stretched lithosphere at first and “hard” collision following the arrival of unstretched Arabian continental lithosphere in the subduction zone. During this evolution, the tectonic regime shows a changeover from extensional (in Eocene) to contractional (in Oligocene-Miocene) regime.

The main Cenozoic magmatic episodes in UDMA are as follows (Förster et al., 1972; Amidi et al., 1984):

1. Effusion of alkaline and intermediate lavas in upper Paleocene-lower Eocene.
2. Deposition of volcano-sedimentary successions by early-middle Eocene phreatomagmatic eruptions.
3. Eruptions of dacitic-rhyolitic pyroclastic rocks and or dacitic domes formed by subaerial eruptions after uplifting the volcano-sedimentary basin during Pyrenean orogeny.
4. Deposition of marine Qom Formation in Oligocene-Miocene times together with alkaline lava flows.
5. Subaerial andesitic eruptions and small intrusions in lower Miocene.
6. Explosive and effusive eruptions with acidic intermediate in composition in upper Miocene-Pliocene.

The magmatic sequence of the studied area can be discriminated into three main facies based on lithostratigraphy and age relations (Figure 2): 1) Eocene volcanics; 2) Sub-volcanic bodies; 3) Miocene lavas.

**Eocene volcanics**

Most volcanic edifices which cropped out in the study area contain the Eocene pyroclastic deposits and volcanic rocks underlain by Cretaceous limestones with an angular unconformity and overlain by Oligocene and Miocene Formations (Emami, 1991). This facies is in turn subdivided into three sub-facies:

*Volcano-sedimentary deposits*

The sedimentary-pyroclastic succession of the Tafresh area is initiated with the Paleocene conglomerate equivalent to Kerman Conglomerate (Huckriede et al., 1962) contained rounded clasts (up to 40 cm in size) and followed by Eocene sedimentary (tuffaceous sandstone, shale, marl and limestone) and green pyroclastic (lithic-crystal tuff and crystal tuff) rocks (Figure 3a). The latter deposits contain benetic microfaunas (such as *Operculina*, *Ditrup* (Figure 3b), *Discocyclina*, *Numulites*) and graded bedding. Because of this, it may be inferred that the shallow-depth sedimentary basin had been interrupted by subaqueous explosive eruptions.

*Andesitic lava flows*

The folded volcano-sedimentary deposits are overlain by andesite (including pyroxene- and amphibole andesite) and basaltic andesite. Because of the occurrences of the reddened soil horizons under lava flows (Figure 3c) and also columnar joints (Figure 3d), it can be stated that the effusive eruptions had been occurred in a subaerial setting after uplifting the basin due to a compression regime.

Petrographically, these lava flows are characterized by porphyritic, megaporphyritic, intersertal and glomeroporphyritic textures and major minerals of plagioclase (andesine-labradorite), diopside, augite, hypersthene, amphibole (magnesian hastingsite and magnesian hornblende) as well as minor minerals such as titanomagnetite and ilmenite. Plagioclase phenocrysts show oscillatory zoning and sieve textures and pyroxene phenocrysts show zoning and exsolution lamellae. Also, many amphibole phenocrysts show evidence of transformations after pyroxenes.

*Thin ignimbrite sheet*

There is a limited and thin (10-20 m in thick) outcrop of rhyolitic ignimbrite with an autaxitic texture and phenocrysts of plagioclase, sanidine, quartz and to a lesser extent, pyroxene and opaque minerals.

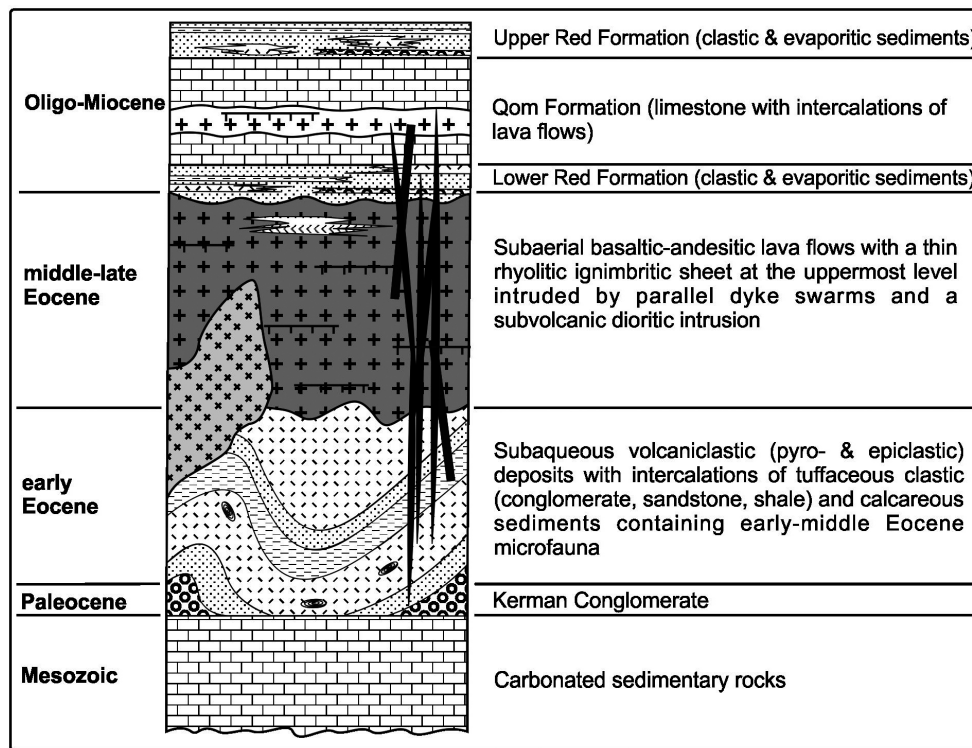


Figure 2. The facies-stratigraphy interactive column for volcano-sedimentary succession in the Tafresh area (without scale).

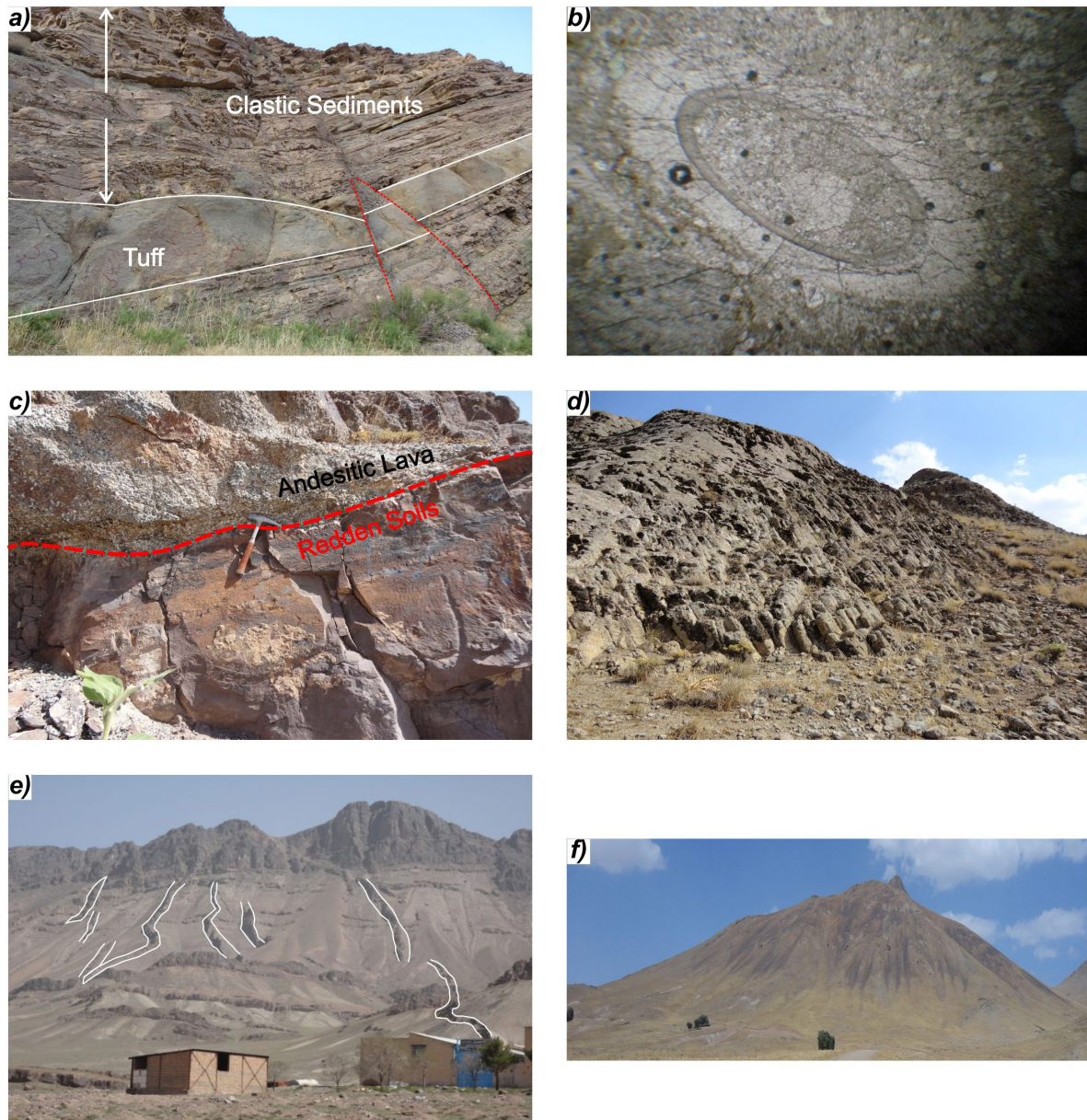


Figure 3. The selected figures from volcanic facies in the Tafresh area: a) intercalations of clastic and pyroclastic deposits; b) *Ditrup* microfossil characterizing Ypresian (early Eocene) in the calcareous sediments; c) the Eocene andesitic lava flow above the marl unit and formation of the reddened soil at their contact; d) columnar joint in the the Eocene andesite lava flow; e) the parallel dyke swarms injected in the succession of lava and pyroclastic deposit succession (NE Tafresh); f) dioritic stock intruded in the Triassic sediments of Zaghar area.

### Subvolcanic bodies

#### Parallel dyke swarms

The parallel dyke swarms are seen in eastern and western sides of Tafresh city (Figure 3e). They have 20-30 m in length and 2-5 m in width and their dominant striking is N40-45W and N10-20W. Because of their localization in the area with the greatest volume of lava flows, it may be regarded them as the feeder dykes through which the andesitic lava flow reached to the surface.

#### Dioritic intrusion

The only intrusive body in the area is a dioritic body that injected as a stock in the Triassic sediments (Figure 3f). The samples show zoned plagioclase and amphibole phenocrysts in a microgranular groundmass.

#### Miocene lavas

These andesitic lava flows have limited outcrops in UDMA (e.g., Bolourchi, 1979; Omrani et al., 2008).

In the study area, a 20-50m lava flow interlayer is seen in the Miocene Qom Formation (Figure 2). Petrographically, it is composed of zoned plagioclase, augite, amphibole (tschermakite and magnesiohastingsite) and titanomagnetite phenocrysts.

### GEOCHEMISTRY

For geochemical whole-rock considerations, 9 fresh samples of various rock types were selected and analyzed by XRF and ICP-OES methods (Table 1). Samples were crushed and pulverized to 60-70 $\mu$ m. Whole-rock major elements were determined by XRF spectrometry (Philips MagiXPRO pw2540) after the preparation as the pressed pellets containing 4 grams of each sample and 0.8 gr Hoechst Wax-C after homogenizing in an agate mortar. The trace and rare earth elements were analyzed by ICP-OES method (Varian-735Es). 0.5 g of each sample was mixed by 10 ml hydrofluoric acid and 3 ml of perchloric acid and finally was heated up to 160 °C. The volume of solution should be about 1-2 ml which after cooling, 2 ml nitric acid and 3 ml of hydrochloric acid would be added to it to gain a clear solution for analysis.

For mineralogical surveying, 2 samples (28 points) in RomaTre University (Italy) and 6 samples (343 points) in University of Adelaide (Australia) were analysed by electron microprobe method. Mineral analyses were obtained with a CAMECA SX51 (15 KV, 20 nÅ) electron microprobe at the University of Adelaide, Australia and with a CAMECA SX50 (15 KV, 15 nÅ) electron microprobe at the RomaTre University, Italy.

### Whole rock geochemistry

As shown in Table 1 and Figure 4, the samples are belonged to subalkaline suite with a calc-alkaline affinity ( $\text{SiO}_2=54.4-72.4$ ;  $\text{Na}_2\text{O}+\text{K}_2\text{O}=2.8-7.7$ ;  $\text{Na}_2\text{O}-2<\text{K}_2\text{O}$ ). Based on the chondrite-normalized REE (Figure 5a) and primitive mantle-normalized multi-element spiderdiagrams (Figure 5b), they show many characteristics of subduction zone magmas including: LREEs enrichment relative to HREEs, negative anomaly of Nb and Ti and also strong positive anomaly of Pb. However, they don't show any negative anomaly of Ta, as a distinctive geochemical signature of subduction zone magmatism. Enrichment of LILEs and positive anomaly of Pb and U could be produced by fluids emanated from subducting lithosphere into the mantle wedge and addition of plagic sediments into the melting source, respectively (Rollinson, 1993; Tatsumi et al., 2002; Varekamp, 2010). Also, the subcontinental lithospheric mantle can be envisaged as another alternative source for these enrichments.

On the other hand, most samples occupy the field intercontinental rift of Cabanis and Lecolle (1989) diagram

(Figure 6a) and within plate volcanic zone (WPVZ) field of Gorton and Schandl (2000) diagram (Figure 6b) and lie near the OIB magmatic source (Figure 6c). Kumar et al. (2015) suggest that La/Nb ratio is a useful geochemical proxy for distinguishing melts derived in subduction-zone and rift-zone settings. Accordingly, the studied samples set in the rift-related magmas (Figure 6d). Also, they set in the transitional region between the island arc and within-plate fields of Condie (1989) and D'Orazio et al. (2004) plots (Figures 6 e,f). Thus, we propose a subduction-related magma source in an intercontinental extensional regime.

Also, because of high amount of Th (7.49-16.69), high ratio of Ba/La (>15) and positive anomaly of Pb, it can be stated that a crustal component together with oceanic sediments had been played a considerable role in the chemical signatures of magma sources in the studied area (Wood, 1980; Fan et al., 2003).

### Mineral chemistry

#### Feldspars

Based on the Figure 7, the composition of feldspars (Table 2) in the samples shows the following characteristics:

- i) The phenocrysts and microliths in the Eocene lavas are labradorite ( $\text{An}_{47-72}$ ) and Na-K feldspars ( $\text{An}_{17-37}$  and sanidine) in composition, respectively.
- ii) The feldspars compositions in the dioritic intrusion show the same fields of Eocene lavas ( $\text{An}_{51-56}$  and  $\text{An}_{16-28}$ ).
- iii)  $\text{An}_{46-70}$  and K-feldspar (mostly anorthoclase) are the composition of feldspars in the Miocene lavas.

#### Pyroxenes

Whereas the pyroxene phenocrysts (Table 3) in the Miocene lavas are diopside in composition, the ones in the Eocene lavas set in two types (Figure 8a):

- i) diopside ( $\text{Fe}\#<0.2$ , low in Ti and high in Cr and Ca)
- ii) pigeonite ( $\text{Fe}\#>0.35$ , high in Ti and low in Cr and Ca).

Furthermore, many pyroxene phenocrysts of Eocene lavas show the exsolved lamellae and fall in the volcanic arc and within plate fields of Nisbet and Pearce (1977) and Leterrier et al. (1982) (Figure 8b).

#### Amphiboles

All amphiboles in the studied samples (Table 4) have  $\text{mg}\#>0.5$ ,  $\text{Ca}_B>0.5$  (1.23-1.88) and  $\text{Ti}<0.5$  (0.046-0.26). But, as shown in Figure 9, they are subdivided into two types:

- i) magnesiohornblende ( $\text{mg}\#=0.76-1$ ,  $(\text{Na}+\text{K})_B=0.27-0.48$ ) in the diopside-bearing Eocene lavas and dioritic intrusion.
- ii) magnesiohastingsite ( $\text{mg}\#=0.73-1$ ,  $(\text{Na}+\text{K})_B=0.4-$

Table 1. Major element (wt%) and trace element (ppm) data for volcanic rocks from the Tafresh area, central Iran.

Sample No.	14/F	4/E-1	11/A	9/E-2	11/F	10/F-6	1/F	6/E	C-1
	Eocene Lavas		Dyke Swarms		Dioritic Intrusion		Ignimbrite Sheet		Miocene Lavas
SiO <sub>2</sub>	54.40	51.40	66.30	62.80	56.90	60.80	72.40	70.30	64.80
TiO <sub>2</sub>	0.70	1.20	0.90	1.00	0.70	0.50	0.40	0.60	0.60
Al <sub>2</sub> O <sub>3</sub>	17.40	18.50	14.10	15.10	18.30	17.50	14.30	15.10	15.40
Fe <sub>2</sub> O <sub>3</sub>	9.80	9.50	6.90	7.80	7.10	6.30	1.50	3.30	5.40
MnO	0.20	0.10	0.10	0.10	0.10	0.20			0.10
MgO	4.20	2.70	2.40	2.80	3.00	2.80	0.30	0.80	1.50
CaO	8.20	7.30	0.80	1.50	6.80	5.20	1.30	0.80	5.90
Na <sub>2</sub> O	2.20	2.70	3.90	4.00	2.60	2.70	1.90	2.80	2.40
K <sub>2</sub> O	0.60	1.50	1.10	1.30	0.80	1.20	5.80	4.00	1.20
P <sub>2</sub> O <sub>5</sub>	0.10	0.20	0.20	0.40	0.20	0.20		0.10	0.30
LOI	1.89	4.66	3.08	3.00	3.24	2.60	1.93	2.05	1.95
Total	99.79	99.76	99.78	99.80	99.74	100.00	99.83	99.85	99.85
Cs	5.72	11.21	8.35	8.48	9.15	5.20	1.22	3.08	4.81
Rb	81.07	48.52	13.78	22.66	78.11	106	11.67	22.46	136
Ba	175	256	220	153	409	507	516	702	510
Sr	332	198	56	88	312	432	47	89	517
Pb	24.63	32.68	26.11	21.08	13.48	21.70	8.53	8.01	20.63
Th	15.73	15.95	12.53	12.25	14.14	9.56	12.97	7.49	16.69
U	4.60	8.96	6.17	6.28	7.42	4.24	0.54	2.40	3.96
Zr	238	344	286	332	247	122	68	392	331
Hf		8.65	1.73	6.27	9.25	6.98	8.55	10.46	13.38
Ta	4.99	7.96	3.20	2.00	2.28	2.25	0.88	0.99	1.99
Y	20.43	32.23	24.62	24.54	21.59	22.84	23.84	25.35	19.34
Nb	25.18	43.42	29.37	34.24	28.40	23.40	12.16	19.91	23.78
Sc	28.63	25.23	17.51	15.54	12.81	13.30	6.35	8.20	12.50
V	218	234	52	70	155	106	19	27	131
Cr	101.72	16.12	19.78	16.74	18.53	28.36	19.64	73.12	17.67
Co	20.33	28.91	16.78	19.35	22.88	13.86	1.15	6.11	10.09
Ni	36.97	11.39		7.64	4.40	1.96	2.17	11.90	
Cu	32.32	44.60	9.37	5.42	22.33	18.86	9.42	3.92	30.94
Zn	69.35	152.51	168.16	222.85	140.94	78.77	16.76	50.65	52.57
Ga	18.27	20.24	17.97	19.41	24.80	13.87	8.84	15.40	14.29
Ge	1.19	3.01	2.17	2.09	2.31	1.22	<0.4	1.02	1.37
As	4.53	17.59	0.50	4.00	6.04	2.62	3.30	12.88	9.82
Se		0.65	0.57			0.30			0.61
Ag		0.10	0.16	0.10	0.20			0.10	0.16
Cd						0.41	0.89		0.58
In	0.93	0.55			0.67	0.83			
Sn	4.49	4.07	3.12	3.46	3.75	3.14	0.82	1.44	3.44
W	1.09	1.83	1.30	1.67	1.59	0.91	0.77	1.01	1.12
La	8.50	16.06	11.65	10.96	14.46	20.71	30.85	19.67	20.85

Table 1. Continued ...

Sample No.	14/F	4/E-1	11/A	9/E-2	11/F	10/F-6	1/F	6/E	C-1
	Eocene Lavas		Dyke Swarms		Dioritic Intrusion		Ignimbrite Sheet		Miocene Lavas
Ce	22.43	41.75	21.93	22.16	34.66	42.46	64.29	37.49	41.36
Pr	6.68	7.40	8.49	6.76	6.07	9.26	14.73	6.31	10.00
Nd	11.01	21.54	7.31	7.48	11.76	21.19	32.07	19.01	13.35
Sm	1.73	4.20	2.32	2.64	5.07	4.25	6.09	3.69	4.38
Eu	1.07	1.89	0.80	0.92	1.14	1.17	0.44	1.09	1.28
Gd	5.82	6.43	4.59	5.22	5.36	5.12	2.78	3.95	4.94
Tb	1.07	0.98	0.78	0.81	0.85	0.76	0.24	0.43	0.74
Dy		8.57	5.72	4.14	1.53	1.33	9.55	8.98	2.56
Ho	0.99	0.91	0.58	0.61	1.03	0.90	0.49	0.42	0.93
Er	2.57	3.21	2.11	2.62	2.51	2.24	1.71	2.04	1.14
Tm	0.51	0.49	0.33	0.37	0.42	0.32		0.15	0.35
Yb	2.74	3.71	2.84	3.05	2.60	2.32	2.19	2.64	2.38
Lu	0.56	0.51	0.34	0.43	0.46	0.38		0.17	0.39

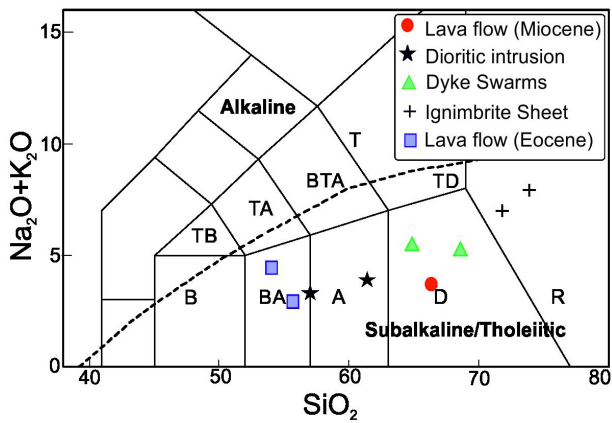


Figure 4. TAS classification diagram of Le Maitre et al. (2002). Abbreviations: A: andesite; B: basalt; BA: basaltic andesite; BTA: basaltic trachyandesite; D: dacite; R: rhyolite; T: trachyte; TA: trachyandesite; TB: trachybasalt; TD: trachydacite.

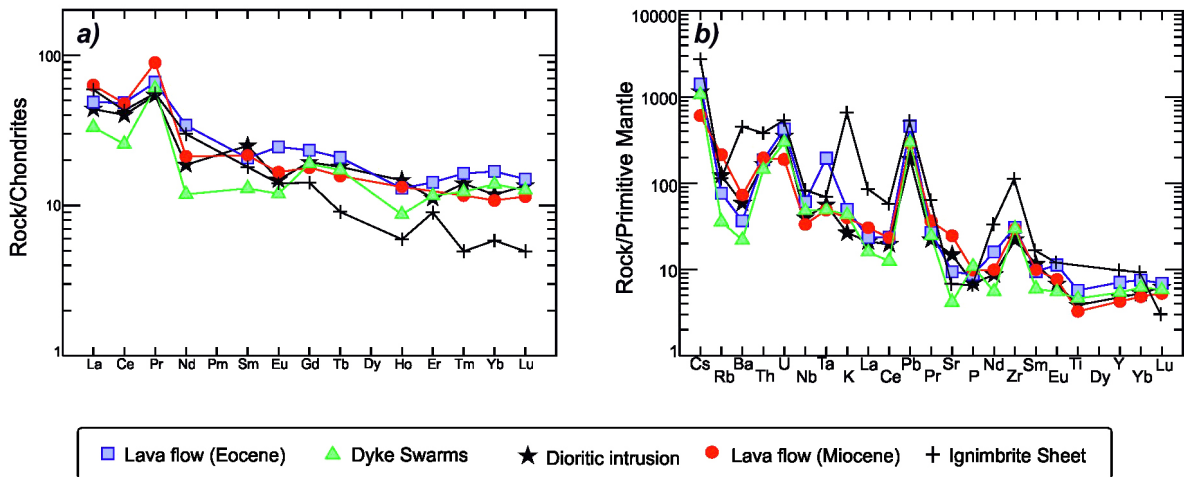


Figure 5. a) REE-spiderdiagram normalized by chondrite (Nakamura, 1974); b) multi-element spiderdiagram normalized by primitive mantle (Sun and McDonough, 1989).

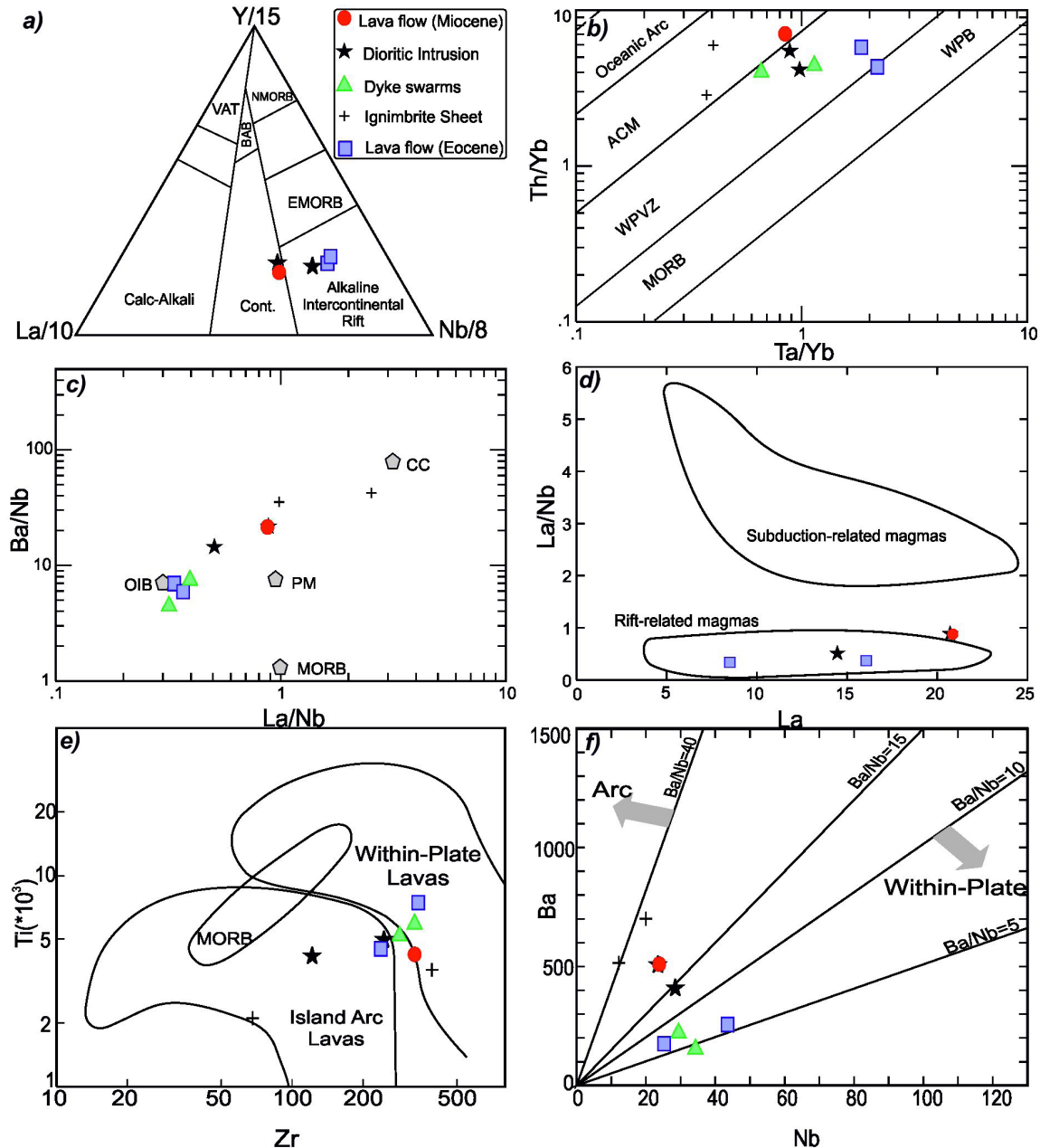


Figure 6. Tectonic discrimination plots: a) La/10-Y/15-Nb/8 plot (Cabanis and Lecolle, 1989); b) Th/Yb vs Ta/Yb diagram (Pearce, 1983 with changes from Gorton and Schandl, 2000) (ACM =Active Continental Margins; WPB =Within Plate Basalts; WPVZ =Within Plate Volcanic Zones); c) Ba/Nb vs. La/Nb plot (PM = primary mantle; OIB = ocean island basalt; MORB =mid-ocean ridge basalt; CC =continental crust) (data of PM, OIB, and MORB are from Sun and McDonough (1989), whereas CC data are from Rudnick and Gao (2003)); d) La/Nb vs. La plot (Kumar et al., 2015); e) Ti/Zr diagram of Condie (1989); f) Ba/Nb diagram of D’Orazio et al. (2004).

0.73,  $^{VI}Al < Fe^{+3}$ ) in the pigeonite-bearing Eocene lavas and Miocene lavas.

So, we can discriminate the Eocene lavas into two mineralogical types: 1) diopside-magnesiohornblende-bearing andesites (DHA) and 2) pigeonite-magnesiohastingsite-bearing andesites (PHA).

#### Oxygen fugacity

According to Spear (1981) and Anderson and Smith (1995), oxygen fugacity of magmas can be calculated from  $Fe\#$  vs  $Al^{IV}$  in amphiboles. Accordingly, as shown in Figure 10, all samples fall in the high oxygen fugacity field. As seen in Figure 10, the PHA show the higher



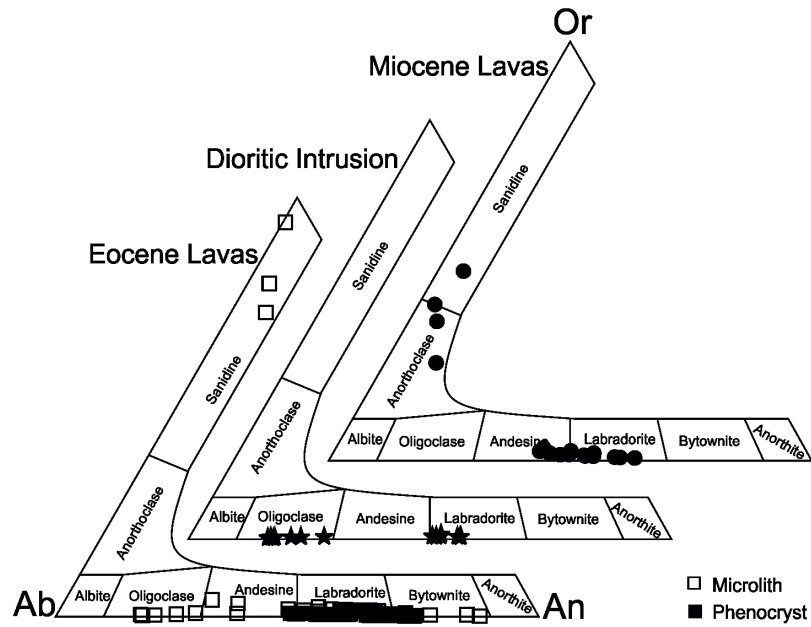


Figure 7. Chemical compositions of feldspars in the volcanic rocks of the study area plotted in an Or-Ab-An ternary diagram.

Table 2. Representative chemical compositions and calculated mineral formulae and modal mineral composition of feldspar.

	Eocene lavas						Dioritic intrusion				Miocene lava			
	Pl		Kfs				Pl		Kfs		Pl		Kfs	
Sample	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	10/F-6	10/F-6	10/F-6	10/F-6	C-1	C-1	C-1	C-1
SiO <sub>2</sub>	54.59	49.46	64.17	65.47	65.08	68.22	54.88	53.32	64.27	65.53	55.89	52.09	66.34	67.91
Al <sub>2</sub> O <sub>3</sub>	26.64	31.46	20.38	19.39	23.57	23.85	27.38	28.13	22.60	23.00	27.12	29.48	20.91	19.64
Cr <sub>2</sub> O <sub>3</sub>	0.01	0.03	0.03	0.00	0.01	0.03	0.01	-0.02	-0.02	0.00	-0.02	0.00	-0.02	0.01
FeO	0.23	0.29	0.05	0.09	0.05	0.07	0.23	0.26	0.05	0.03	0.59	0.51	0.62	0.63
CaO	10.05	15.13	0.80	0.08	3.50	3.66	10.62	11.41	3.60	3.30	10.05	12.80	2.15	0.81
Na <sub>2</sub> O	5.78	2.91	1.61	0.52	9.21	3.38	5.59	4.91	9.80	6.81	5.64	4.19	7.48	4.21
K <sub>2</sub> O	0.14	0.05	12.09	13.80	0.11	0.07	0.13	0.11	0.09	0.07	0.39	0.10	3.96	5.73
Total	98.90	99.33	100.32	99.87	101.49	99.30	98.90	98.22	100.37	98.72	99.76	99.24	102.02	99.20
No. Oxygens	8	8	8	8	8	8	8	8	8	8	8	8	8	8
Si	2.52	2.28	2.93	3.00	2.82	2.93	2.50	2.46	2.83	2.88	2.53	2.39	2.91	3.02
Al	1.45	1.71	1.10	1.05	1.20	1.21	1.47	1.53	1.17	1.19	1.45	1.59	1.08	1.03
Fe <sup>2+</sup>	0.01	0.01	0.00	0.00	0.00	0.00	0.01	0.01	0.00	0.00	0.02	0.02	0.02	0.02
Ca	0.50	0.75	0.04	0.00	0.16	0.17	0.52	0.56	0.17	0.16	0.49	0.63	0.10	0.04
Na	0.52	0.26	0.14	0.05	0.77	0.28	0.49	0.44	0.84	0.58	0.49	0.37	0.64	0.36
K	0.01	0.00	0.70	0.81	0.01	0.00	0.01	0.01	0.01	0.00	0.02	0.01	0.22	0.33
Total	5.01	5.00	4.92	4.90	4.97	4.60	5.01	5.00	5.01	4.81	5.00	5.01	4.97	4.80
%Or	0.80	0.27	79.52	94.11	0.63	0.87	0.76	0.62	0.51	0.54	2.26	0.60	23.09	44.72
%Ab	50.56	25.74	16.08	5.44	82.11	62.02	48.42	43.52	82.69	78.48	49.24	37.00	66.36	49.95
%An	48.63	73.98	4.40	0.45	17.27	37.11	50.82	55.86	16.80	20.98	48.50	62.40	10.55	5.33

Table 3. Representative chemical compositions and calculated mineral formulae and modal mineral composition of clinopyroxene.

Sample	Eocene lavas										Miocene lava		
	Cpx					Opx					Cpx		
	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	14/F-1	c-1	c-1
SiO <sub>2</sub>	53.41	51.74	51.07	51.27	50.69	51.12	51.43	51.40	51.67	51.72	51.74	51.01	49.06
TiO <sub>2</sub>	0.20	0.30	0.58	0.61	0.67	0.37	0.36	0.27	0.37	0.41	0.28	0.40	0.19
Al <sub>2</sub> O <sub>3</sub>	1.92	2.35	1.47	1.49	1.35	3.22	1.92	0.99	0.93	1.26	2.02	2.66	0.98
Cr <sub>2</sub> O <sub>3</sub>	0.12	0.19	0.00	0.02	-0.03	0.22	0.05	0.00	0.02	-0.02	0.00	0.00	0.00
FeO	5.33	7.05	14.64	14.37	16.94	8.12	20.03	19.03	20.23	19.16	9.33	8.47	9.34
MnO	0.18	0.18	0.39	0.37	0.48	0.29	0.52	0.63	0.57	0.63	0.47	0.27	1.13
MgO	17.99	16.63	14.38	14.33	12.24	16.59	18.19	18.29	18.62	16.97	15.36	15.05	15.08
CaO	20.77	19.88	15.69	16.04	16.18	18.42	6.05	7.34	5.67	8.64	20.12	21.71	20.93
Na <sub>2</sub> O	0.24	0.21	0.24	0.25	0.27	0.27	0.44	0.12	0.09	0.16	0.25	0.28	0.46
K <sub>2</sub> O	0.01	0.01	0.00	0.00	0.01	0.00	0.02	0.02	0.01	0.02	0.013	0.00	0.05
Total	100.17	98.54	98.47	98.73	98.80	98.73	99.08	98.11	98.11	98.95	99.58	99.84	97.23
No. Oxygens	6	6	6	6	6	6	6	6	6	6	6	6	6
Si	1.95	1.93	1.96	1.96	1.96	1.91	1.95	1.97	1.98	1.97	1.93	1.89	1.88
Ti	0.01	0.01	0.02	0.02	0.02	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01
Al	0.08	0.10	0.07	0.07	0.06	0.14	0.09	0.04	0.04	0.06	0.09	0.12	0.04
Fe <sup>2+</sup>	0.16	0.22	0.47	0.46	0.55	0.25	0.64	0.61	0.65	0.61	0.28	0.25	0.28
Mn <sup>2+</sup>	0.01	0.01	0.01	0.01	0.02	0.01	0.02	0.02	0.02	0.02	0.01	0.01	0.04
Mg	0.98	0.93	0.82	0.82	0.71	0.92	1.03	1.04	1.06	0.96	0.85	0.83	0.86
Ca	0.81	0.80	0.64	0.66	0.67	0.74	0.25	0.30	0.23	0.35	0.80	0.86	0.86
Na	0.02	0.02	0.02	0.02	0.02	0.02	0.03	0.01	0.01	0.01	0.02	0.02	0.03
Total	4.01	4.01	4.00	4.00	4.00	4.01	4.01	4.00	4.00	4.00	4.00	4.00	4.00
%Wo	41.56	40.96	33.29	33.98	34.85	38.50	12.86	15.41	11.97	18.30	41.40	44.31	43.06
%En	50.12	47.70	42.46	42.26	36.68	48.25	53.86	53.42	54.70	50.03	43.95	42.72	43.17
%Fs	8.32	11.34	24.25	23.76	28.48	13.25	33.27	31.17	33.33	31.67	14.65	12.97	13.77

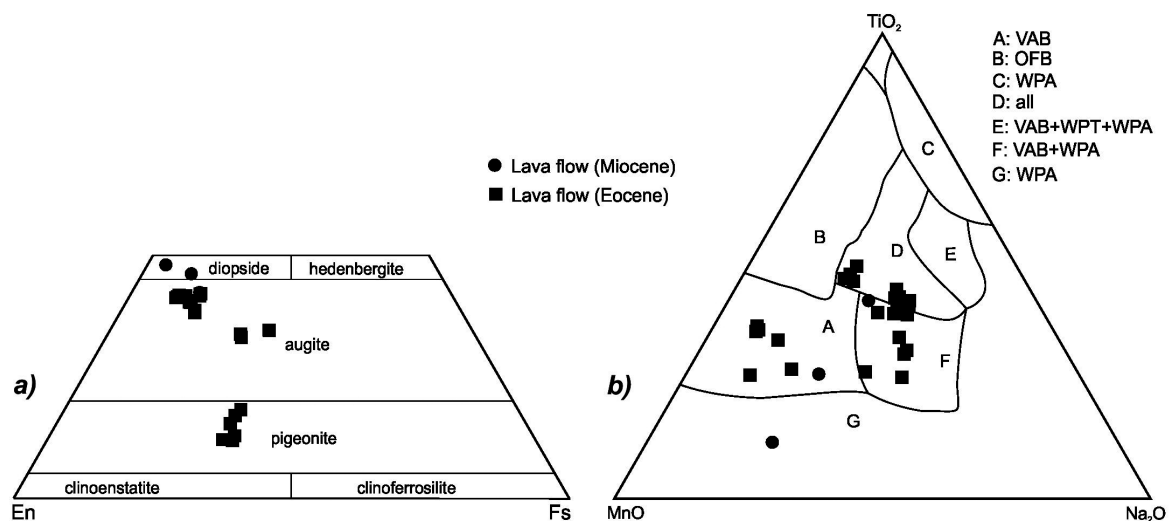


Figure 8. Chemical compositions of pyroxene phenocrysts plotted in: a) an En-Wo-Fs ternary diagram (Morimoto et al., 1988); b) a tectonic classification (Nisbet and Pearce, 1977).

Table 4. Representative chemical compositions and calculated mineral formulae of amphibole.

Sample	Dioritic intrusion				Eocene lavas				Miocene lava			
	10/F-6	10/F-6	10/F-6	10/F-6	14/F-1	14/F-1	14/F-1	14/F-1	C	C	C	C
	Am											
SiO <sub>2</sub>	45.67	45.00	47.58	46.94	46.72	49.23	47.96	45.72	42.39	42.88	42.49	40.99
TiO <sub>2</sub>	1.61	1.95	1.26	1.27	1.92	1.15	1.06	2.39	2.13	2.12	2.11	1.93
Al <sub>2</sub> O <sub>3</sub>	8.11	8.51	6.63	6.85	8.08	5.74	7.78	8.18	11.81	11.96	12.01	12.41
FeO	14.22	14.41	13.34	13.48	13.80	14.72	13.97	14.31	11.58	11.57	11.61	12.48
MgO	13.47	13.31	14.91	14.60	12.98	14.82	14.53	13.02	14.52	14.56	14.60	0.25
MnO	0.41	0.34	0.50	0.55	0.37	0.52	0.60	0.37	0.23	0.24	0.22	13.49
CaO	10.59	10.68	10.36	10.34	10.39	10.39	10.08	10.88	11.43	11.47	11.50	11.27
Na <sub>2</sub> O	1.35	1.58	1.30	1.26	1.56	1.22	1.29	1.77	2.32	2.28	2.31	2.48
K <sub>2</sub> O	0.20	0.23	0.15	0.15	0.22	0.15	0.12	0.25	0.40	0.39	0.38	0.44
Total	95.63	96.01	96.03	95.44	96.04	97.94	97.39	96.89	96.81	97.47	97.23	95.74
No. Oxygens	23	23	23	23	23	23	23	23	23	23	23	23
Si	6.86	6.75	7.05	7.02	6.88	7.10	6.93	6.72	6.19	6.21	6.17	6.60
Ti	0.18	0.22	0.14	0.14	0.21	0.12	0.11	0.26	0.23	0.23	0.23	0.23
Al	1.43	1.50	1.16	1.21	1.40	0.97	1.32	1.42	2.03	2.04	2.05	2.36
Al <sup>IV</sup>	1.24	1.35	1.03	1.08	1.12	0.89	1.07	1.28	1.81	1.79	1.82	1.39
Al <sup>VI</sup>	0.17	0.12	0.10	0.11	0.28	0.08	0.25	0.13	0.22	0.25	0.23	0.97
Fe <sup>2+</sup>	1.78	1.81	1.65	1.68	1.72	1.80	1.71	1.78	23.00	23.00	23.00	23.00
Mn <sup>2+</sup>	0.05	0.04	0.06	0.07	0.05	0.06	0.07	0.05	0.03	0.03	0.03	1.84
Mg	3.01	2.98	3.30	3.25	2.88	3.22	3.17	2.88	6.28	6.31	6.27	6.11
Ca	1.70	1.72	1.64	1.66	1.66	1.62	1.58	1.73	2.06	2.07	2.09	1.89
Na	0.39	0.46	0.37	0.36	0.45	0.35	0.37	0.51	0.50	0.48	0.50	0.51
K	0.04	0.04	0.03	0.03	0.04	0.03	0.02	0.05	1.43	1.42	1.43	0.70
Total Cations (S)	15.46	15.52	15.41	15.42	15.37	15.38	15.39	15.49	0.08	0.07	0.07	0.08

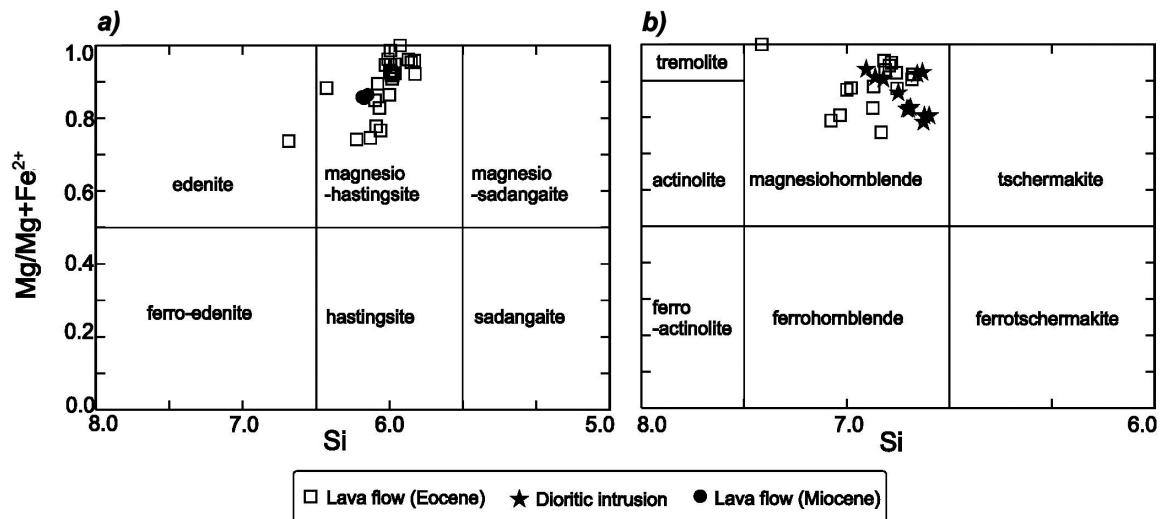


Figure 9. Chemical compositions of amphiboles (Leake et al., 1997): a) Eocene pigeonite-magnesiohastingsite bearing andesites (PHA) and Miocene lavas; b) Eocene diopside-magnesiohornblende bearing andesites (DHA) and dioritic stock.

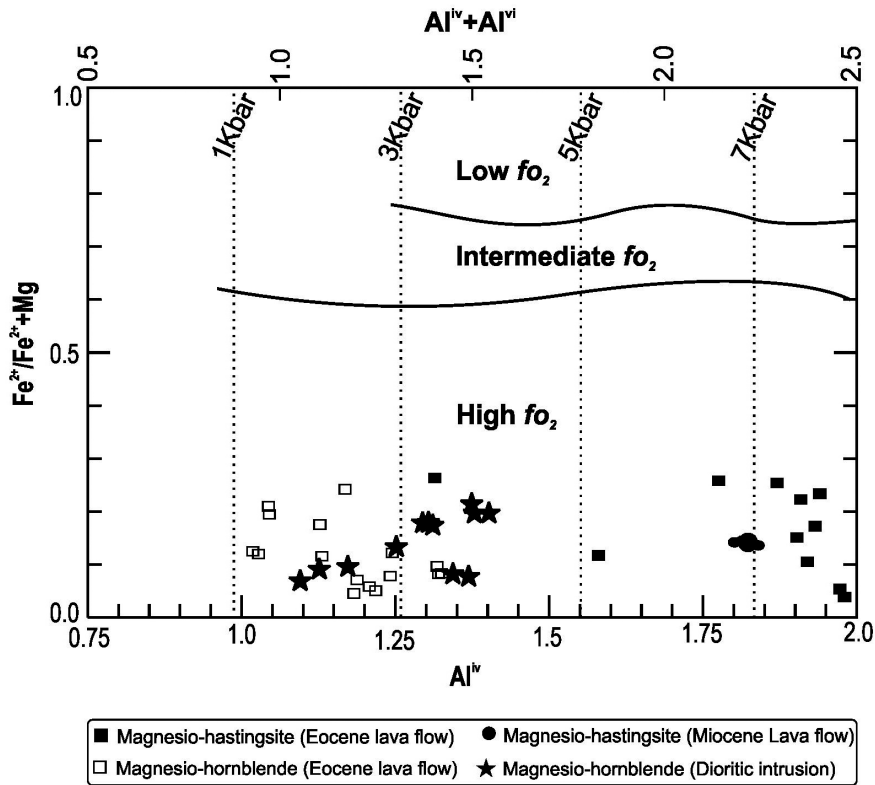


Figure 10. Fe# vs Al<sup>IV</sup> plot for distinguishing of fugacities in amphiboles (Anderson and Smith, 1995) and Fe# vs Al<sup>IV</sup>+Al<sup>VI</sup> for defining of pressures in amphiboles (Smith, 1992).

Table 5. Results of geothermobarometric calculation for different volcanic facies of the Tafresh area.

Rock Unit	P (Kbar)		T (°C)	
Eocene Lavas	DHA*	1-4	Smidth (1992) Helz (1973)	620±20 Elkins and Grove (1990) Putirka (2008)
	PHA**	> 6		900±50 Lindsley (1983)
Dioritic Intrusion		2.5-4.5	Schmidt (1992)	685±10 Holland and Blundy (1994)
		2.5-4.2	Anderson and Smith (1995)	620±20 Elkins and Grove (1990) Putirka (2008)
		2.5-4.0	Ridolfi (2010)	
Miocene Lavas		6.5-8.5	Schmidt (1992)	720±10 Holland and Blundy (1994)
		5.8-8	Anderson and Smith (1995)	875±25 Ernst and Liu (1998)
		5-7	Ridolfi (2010)	1140±45 Elkins and Grove (1990) Putirka (2008)
		7.5	Ernst and Liu (1998)	

\* Diopside-magnesianhornblende bearing andesite  
 \*\* Pigeonite-magnesianhastingsite bearing andesite

pressure (>6 Kbar) than DHA (about 1-4 Kbar) in the Fe# vs <sup>total</sup>Al plot of Schmidt (1992).

### Geothermobarometry

The results of geothermobarometric calculations on different magmatic facies by various methods (i.e., Al-contents in amphiboles (Schmidt, 1992; Anderson and Smith, 1995; Ridolfi et al., 2010), two feldspar thermometer (Elkins and Grove, 1990; Putirka 2008), hornblende-plagioclase pair thermometry (Holland and Blundy 1994) and two-pyroxene thermometry (Lindsley, 1983) have been given in Table 5. As shown in Table 5, Miocene lavas had been formed in higher *P* and *T* than Eocene lavas and dioritic intrusion. Thus, it can be postulated that the Miocene lavas had likely been originated from more depths.

### AFC modeling

AFC modeling was based on the calculations of De Paolo (1981) using the sample 9/E-2 as the parent magma and the continental crust from Taylor and McLennan (1985, 1995) as the contaminant. As seen in the Ce/Pb vs Ba/Nb plot (Figure 11), the samples have a good fitness with AFC curves with  $r=0.7$ , but different *D*-values. Thus, it seems that the magma chamber had been open repeatedly for new pulses of magma and crustal components. Such a conclusion is verified by different mineral paragenesis of feldspars, pyroxenes and amphiboles (Figures 7-9).

### STRUCTURAL ANALYSIS

The main structural elements in the study area recognizable on the satellite and aerial images as well as on the outcrops (Figure 12) are dominantly: i) reverse (thrust) faults with NW-SE strike-slip component; ii) dextral strike-slip faults with NNW-SSE striking; iii) folding with NW-SE striking associated with strike-slip faults. It seems that the strike of the folds has bent bearing of the fault (Figure 12).

According to many researchers (e.g., Nogol-Sadat, 1985; Walker and Jackson, 2004; Alavi, 2007; Morley et al., 2009), the main faults in UDMA, including Qom-Zefreh fault, Bidhend fault Dehshir-Baft fault, are basement faults. Nogol-Sadat (1985) and Mohajjel and Ferguson (2000) propose that the dextral transpressional system in UDMA activated the basement faults. On the other hand, the releasing and restraining zone could be produced by en-echelon pattern of dextral strike-slip faults in UDMA (Babaahmadi et al., 2010). Moreover, there are some overlaps between the compressional terminations of the dextral faults in the area. So, whereas such a structural regime produced the NW-SE striking folds, the parallel dyke swarms have been concentrated on tension gash zone on both sides of Rahjerd fault (Figure 12).

Moreover, Nogol-Sadat (1978) propose the Tafresh, Indes, Talkhab and Chaghar faults as the main dextral faults (N125-130E) and Bidhend fault (N160-170E) as a minor extensional fault in UDMA that was likely had been involved in Cenozoic volcanism. So, because of parallel striking of both Rahjerd and Bidhend faults and also the concentration of parallel dyke swarms on both sides of the Rahjerd fault, it can be stated that the Cenozoic volcanism in the area had been occurred by the extensional role of Rahjerd fault. Also, it seems that the N10-20W dyke swarms were produced by reactivation of Rahjerd fault and its local extension. Accordingly, the latter dyke swarms must be younger than the N40-50W dyke systems.

### DISCUSSION AND RESULTS

Urumieh Dokhtar magmatic arc (UDMA) on the western Iran is a volcano-plutonic zone that proposed by most researchers (e.g., Berberian and King, 1981; Alavi, 1994; Agard et al., 2005; Vincent et al., 2007; Horton et al., 2008; Morely et al., 2009) as a result of subduction of Neothetys oceanic lithosphere and then continental collision of Arabia and central Iran microcontinents. On the other hand, it proposes that the main strike-slip faults (Qom-Zefreh, Dehshir-Baft and Bidhend faults) in this zone played an important role in the magmatic episodes. Because of this, UDMA could be imagined as a fault zone for reaching the magma to the surface.

In the study area located on the middle part of UDMA

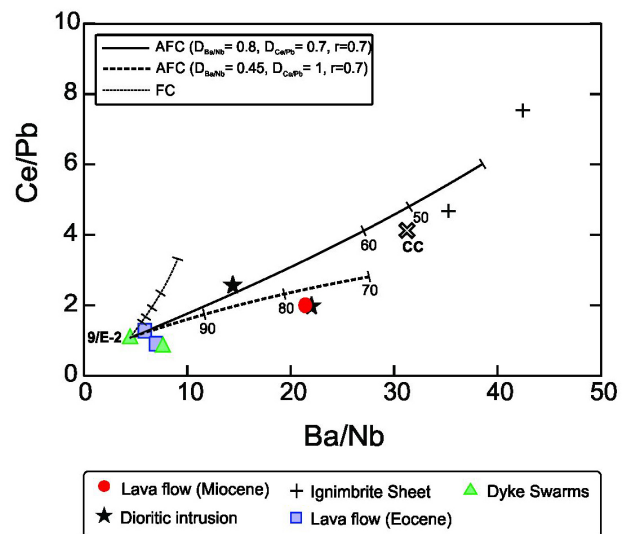


Figure 11. Geochemical modeling for studied samples on Ce/Pb vs Ba/Nb diagram. Note to better fitness of AFC trends than FC trend. The sample No. 9/E-2 have been selected as parental magma (C0) and continental crust (Taylor and McLennan, 1985, 1995) is contaminant (Ca).

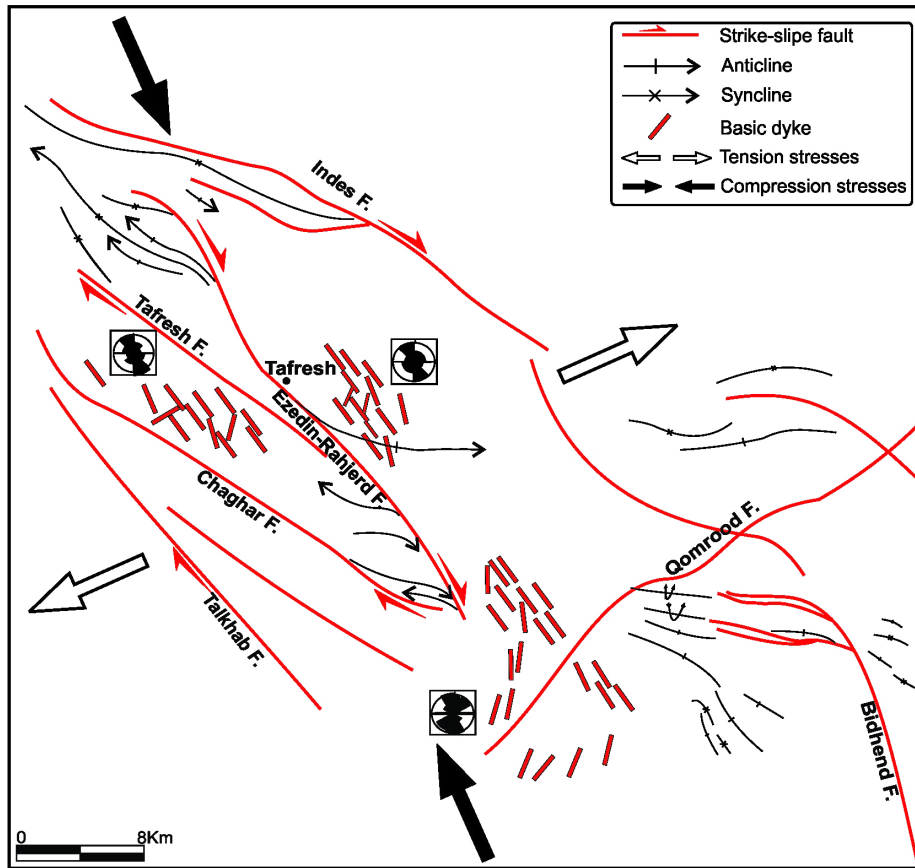


Figure 12. The structural map and the main stress vectors in the studied area. The major shortening trend is based on Hessami et al. (2003).

(near Tafresh city), there are different Cenozoic volcano-sedimentary and volcanic facies. The greatest thickness of lava flows are observed in the places with the parallel dyke swarms. Thus, we call them as feeder dykes. A structural analysis on dyke swarms and faults and folds in the area show that the dyke swarms have been localized on the both sides of the Rahjerd dextral fault (tensional gash zones) (Figure 12). As a result, we suggest that the volcanism in the Tafresh area is a “passive volcanism” during the strike-slip faulting.

Such a tensional regime is verified by lithological changes in the succession in which lower Eocene shallow-depth volcano-sedimentary deposits are followed by the calcalkaline lava flows. According to most researchers in last decades (e.g., Vincent et al., 2005; Ballato et al., 2011; Verdel et al., 2011; Rezaeian et al., 2012), the Eocene magmatism in UDMA was occurred in an extensional regime accompanied by crustal subsidence (Rezaeian et al., 2012). Also, Vincent et al. (2005) suggest that the magmatism and deposition have taken place in an extensional to transtensional setting, possibly caused by roll-back of the subducting Neotethys lithosphere

(Vincent et al., 2005; Verdel et al., 2011). Although, the samples preserve some subduction-related characteristics (such as negative anomaly of Nb and to a lesser extent of Ti, strong positive anomaly of Pb and also enrichment of LILEs), they don't show negative anomaly of Ta (as a distinctive signature of subduction mechanism). Moreover, they fall in the rift-related or within-plate fields of tectonic diagrams of Cabanis and Lecolle (1989) (Figure 6a), Gorton and Schandl (2000) (Figure 6b) and Kumar et al. (2015) (Figure 6d) and or in the transitional region between the island arc and within-plate of Condie (1989) and D'Orazio et al. (2004) plots (Figures 6 e,f). On the other hand, the pyroxene minerals set in both field of VAB and WPA in the diagram of Nisbet and Pearce (1977) (Figure 8b).

Heterogeneous characteristics of the volcanic samples are confirmed by mineral chemistry. It means that the Eocene andesites are subdivided into two subgroups: one having diopside and magnesio-hornblende (DHA) that show shallower pressure (<4Kbar) and other having pigeonite and magnesio-hastingsite (PHA) were originated at greater pressure (>6Kbar). Geothermobarometric calculations

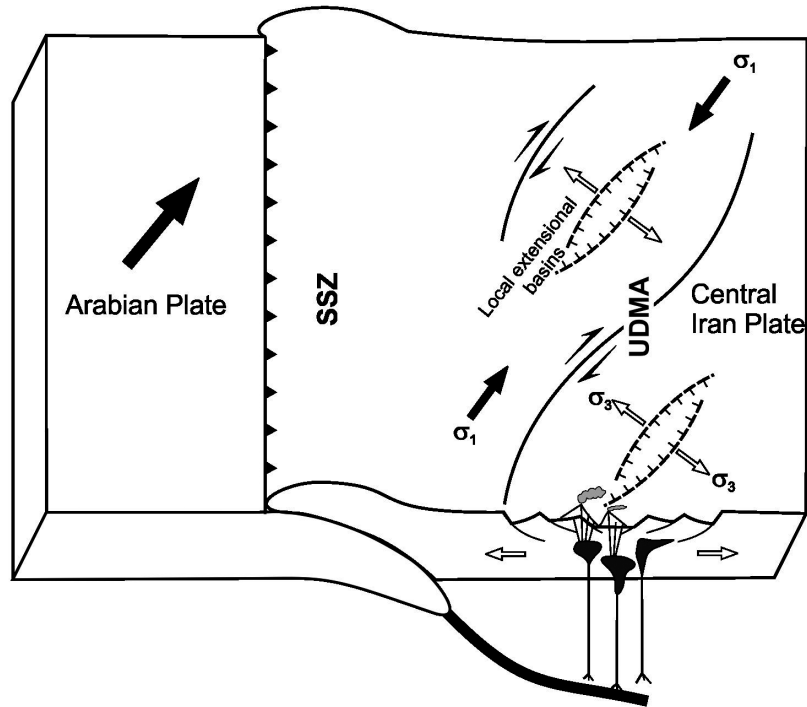


Figure 13. A Schematic model for producing the local tensional basins (tensional gash) between dextral faults in UDMA. Oblique movement trend of Arabian plate is based on Allen et al. (2004).

reveal that DHA and PHA have similar conditions with diorites and Miocene andesites, respectively. On the other hand, the AFC process could be concluded as the dominant process for explaining of magmatic evolution in the area. It can be imagined that the primary magma (with a subduction affinity) had been evolved in the crustal chambers by fractional crystallization and crustal contamination (Figure 13). The dextral fault system in the area (mainly Rahjerd fault) had produced the local extensional basins (Figure 12). The similar process has been reported for intra-arc basins in the Peruvian Andes by Polliand et al. (2005). According to Polliand et al. (2005), dextral wrenching was a trigger for the formation of a series of pull-apart basins and the emplacement of the Coastal Batholith in the Peruvian Andes.

The structural and petrological considerations of this study have given the following results:

1- The volcano-sedimentary succession in the Tafresh area occurred mainly during two consecutive stages: subaqueous and then subaerial stages that produced three major facies: Eocene volcanics (including volcano-sedimentary deposits, andesites and rhyolitic ignimbrites), subvolcanic bodies (including parallel dyke swarms and dioritic intrusions) and Miocene andesites.

2- Geochemically, the magma formed in a subduction- and collision-related setting, was evolved by crustal

contamination processes in the extensional continental environment through the AFC process.

3- The mineral assemblages confirm the unstable conditions in the magma chamber. It means that the Eocene andesites have two different generation of plagioclases ( $An_{47-72}$  and  $An_{17-37}$ ), pyroxenes (diopside and pigeonite) and amphiboles (magnesiohastingsite and magnesiohornblende). Because of this, we have subdivided them into two types: diopside and magnesiohornblende andesites (DHA) and other pigeonite and magnesio-hastingsite andesites (PHA).

4- Geothermobarometric and mineral chemistry considerations reveal that the DHA and dioritic body have similar conditions (1-4 Kbar and 620-850 °C). Whereas, PHA and Miocene lavas show higher pressure and temperatures (>6 Kbar and about 900 °C).

5- The dextral strike-slip fault systems in the area produced the local extensional basin (or tensional gashes) for ascending the magma to the surface through feeder dykes.

6- As a result, we suggest the local extensional basins due to reactivation of NW-SE dextral strike-slip faults for appearance and evolution of different magmatic facies in the Tafresh area.

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